

# Current status of the laser fusion research and the shock ignition approach

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On behalf of the EUROfusion Enabling Research Project ENR-IFE19.CEA-01 “Study of Direct Drive and Shock Ignition for IFE: Theory, Simulations, Experiments, Diagnostics development ” and in particular *J.Trela, G.Boutoux, A.Casner, X.Ribeyre, V.Tikhonchuk, E.Le Bel, Ph.Nicolai, D.Raffestin, A.Colaitis, L.Antonelli, N.Woolsey, G.Cristoforetti, L.Gizzi, R.Scott, K.Glize, S.Pikuz, J.Honrubia, R.Florido, S.Atzeni, T.Pisarczyk, S.Guskov, D.Mancelli, A.Tentori, O.Renner, J.Dostal, R.Dudzak, M.Krus, F.Baffigi, E.Filippov, Y.J.Gu, O.Klimo, S.Malko, A.Martynenko, S.Pikuz, T.Chodukowski, Z.Kalinowska, M.Rosinski, ...*

**Experiments realised in collaboration with the University of Rochester team (Omega laser facility):**

*R.Betti, J.Peebles, W.Theobald, K.S.Anderson, J.A.Delettrez, V.Yu.Glebov, A.A.Solodov, M.Stoeckl, C. Stoeckl, M.Wei, Univ. Rochester, USA*  
*J.A.Frenje, Massachusetts Institute of Technology, Cambridge MA, USA*

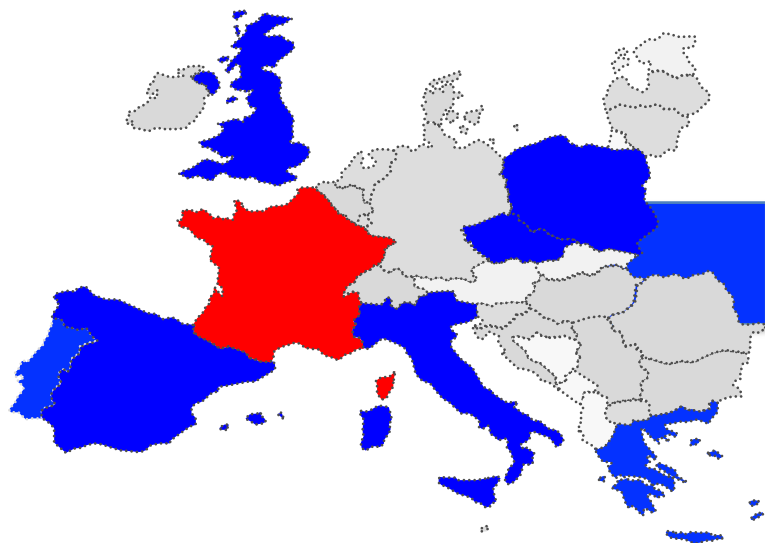


30<sup>th</sup> Summer School and  
International Symposium on  
the Physics of Ionized Gases

August 24 - 28, 2020, Šabac, Serbia



- 1) **CELIA, Bordeaux** - Group Leader: Dimitri Batani
- 2) **Prague team (IoP, ELI, PALS, CTU Prague)** - Group Leader: Miroslav Krous
- 3) **University of Rome "La Sapienza"** - Group Leader: Stefano Atzeni
- 4) **UK shock ignition collaboration (Univ. York, Warwick, Strathclyde, STFC)** - Group Leader: Nigel Woolsey
- 5) **Universidad de Las Palmas de Gran Canaria (ULPGC)** - Group Leader: Ricardo Florido
- 6) **IPPLM, Warsaw** - Group Leader: Tadeusz Pisarczyk
- 7) **INO, CNR, Pisa Italy** - Group Leader: Gabriele Cristoforetti
- 8) **Polytechnic University of Madrid, Spain** - Group Leader: Javier Honrubia
- 9) **CLPU, Salamanca, Spain** - Group Leader: Luca Volpe
- 10) **ENEA, Frascati, Italy** - Group Leader: Fabrizio Consoli
- 11) **IST, Lisbon, Portugal** - Group Leader: Marta Fajardo
- 12) **LULI, Palaiseau, France** - Group Leader: Sophie Baton
- 13) **CPPL of TEI of Crete, Greece** - Group Leader: Michael Tatarakis
- 14) **Kharkov Institute of Physics & Technology, Ukraine** - Group Leader: Vasyl Maslov
- 15) **Università di Milano Bicocca, Italy** - Group Leader: Giuseppe Gorini



**ENR-IFE19-CEA-01 Study of Direct Drive and Shock Ignition for IFE: Theory, Simulations, Experiments, Diagnostics development**

Participants: **FR**, CZ, ES, IT, PL, UK, GR, PT, UKR

# Summary

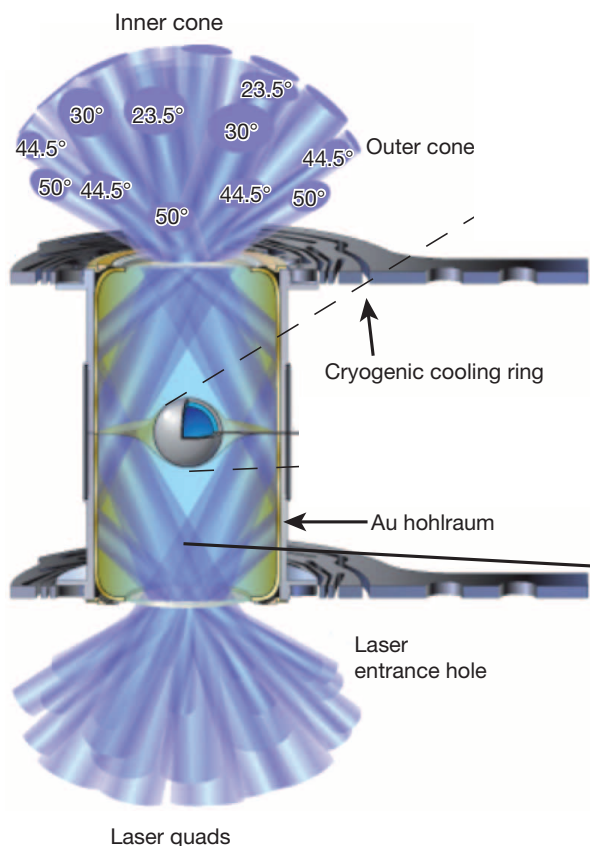
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- 1) What is shock ignition
- 2) The problem of hot electrons and  
Experiments at PALS and Omega-EP
- 3) Roadmap to shock ignition

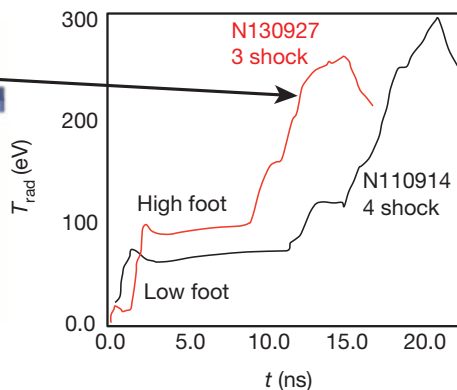
*(many other experiments in several other facilities:  
LIL, LULI, Phelix... Gekko...)*

# NIF:nteresting results after

High-foot implosions (O.Hurricane, et al. Nature 2014) have allowed entering a novel “ $\alpha$ -heating regime”



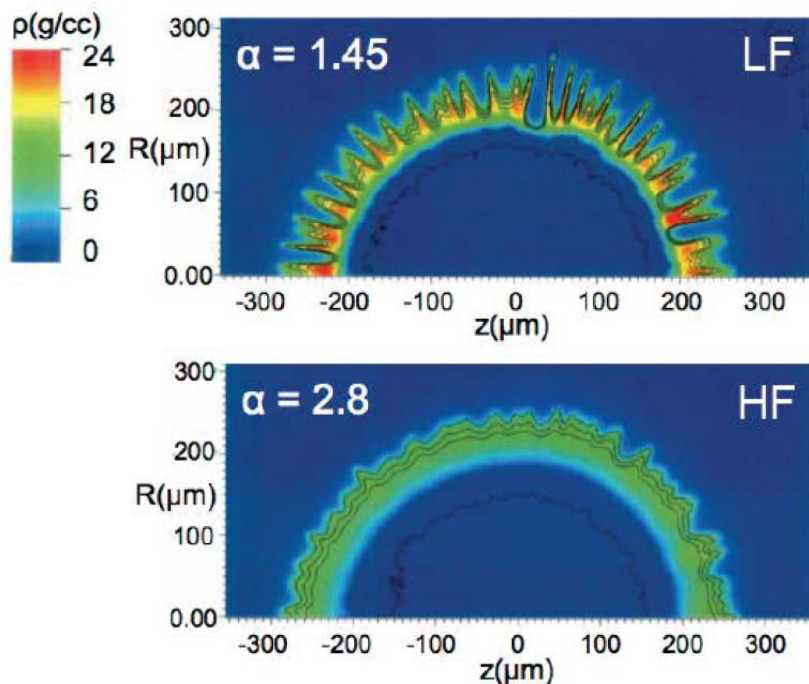
Quantity	N131119 <sup>425 TW</sup> 1.9 MJ	N130927 <sup>390 TW</sup> 1.8 MJ
$Y_{13-15}$ (neutron)	$(5.2 \pm 0.097) \times 10^{15}$	$(4.4 \pm 0.11) \times 10^{15}$
$T_{ion}$ (keV) D-T	$5.0 \pm 0.2$	$4.63 \pm 0.31$
$T_{ion}$ (keV) D-D	$4.3 \pm 0.2$	$3.77 \pm 0.2$
DSR (%)	$4.0 \pm 0.4$	$3.85 \pm 0.41$
$\tau_x$ (ps)	$152.0 \pm 33.0$	$161.0 \pm 33.0$
$PO_x, PO_n$ ( $\mu\text{m}$ )	$35.8 \pm 1.0, 34 \pm 4$	$35.3 \pm 1.1, 32 \pm 4$
$P2/PO_x$	$-0.34 \pm 0.039$	$-0.143 \pm 0.044$
$P3/PO_x$	$0.015 \pm 0.027$	$-0.004 \pm 0.023$
$P4/PO_x$	$-0.009 \pm 0.039$	$-0.05 \pm 0.023$
$Y_{total}$ (neutron)	$6.1 \times 10^{15}$	$5.1 \times 10^{15}$
$E_{fusion}$ (kJ)	17.3	14.4
$r_{hs}$ ( $\mu\text{m}$ )	36.6	35.5
$(\rho)_{hs}$ ( $\text{g cm}^{-2}$ )	0.12-0.15	0.12-0.18
$E_{hs}$ (kJ)	3.9-4.4	3.5-4.2
$E_x$ (kJ)	2.2-2.6	2.0-2.4
$E_{DT,total}$ (kJ)	8.5-9.4	10.2-12.0
$G_{fuel}$	1.8-2.0	1.2-1.4



This also shows that the MAIN problem towards ignition is REALLY the impact of hydro instabilities related to non-uniformities

The best NIF implosions used the High-Foot laser pulse that drives stronger shocks in the “foot”

O. Hurricane, APS DPP meeting (2013)



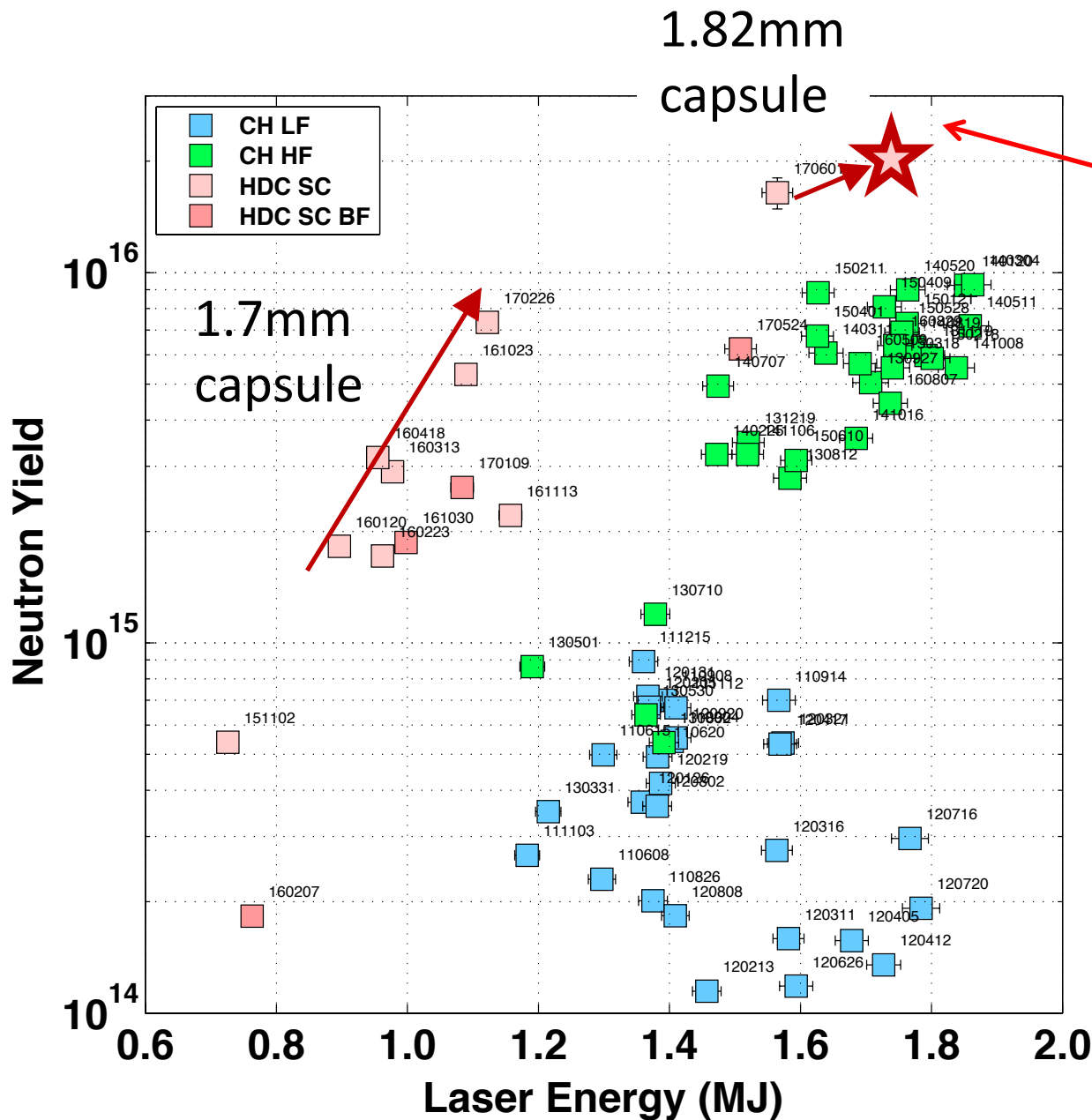
The high foot pulse set the imploding shell on a higher isentrope  $\alpha$  (nothing to do with  $\alpha$ -particles) because it launches stronger shocks in the “foot” of the pulse

$$\gamma = \sqrt{\frac{Akg}{1 + kL}} - \beta kv_{abl}$$

Increase with  $\alpha$

High-foot growth-factor calculations and simulations are consistent with the expectation of less instability

# Very recent results on NIF



**W-doped HDC capsule driven in a low-gas-fill hohlraum 390 km/s, 2e16, 57 kJ of fusion yield, more than 2x  $\alpha$ -heating**

Near Vacuum Hohlraums reduce Laser-Plasma-Instabilities

HDC (diamond) or Beryllium Ablator have greater hydrodynamic efficiency allowing a more massive (and more stable) shell to be imploded

Rugby hohlraum

# Direct Drive ICF

1) The impact of hydrodynamic instabilities (Rayleigh-Taylor) is the main obstacle to achieving ignition in indirect-drive inertial confinement fusion experiments

2) For future reactors we need DIRECT DRIVE:

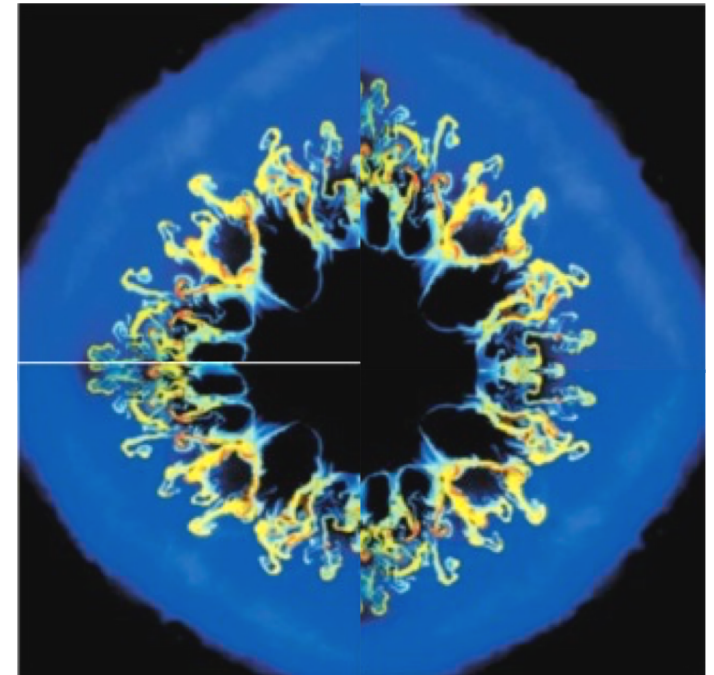
- Higher gains
- Smaller laser facilities
- Simpler targets and simpler scheme more compatible with high-repetition rate operation and requirements of fusion reactors

Unfortunately Direct Drive is even more prone to uniformity problems and hydro-instabilities

**Possible Solution:**

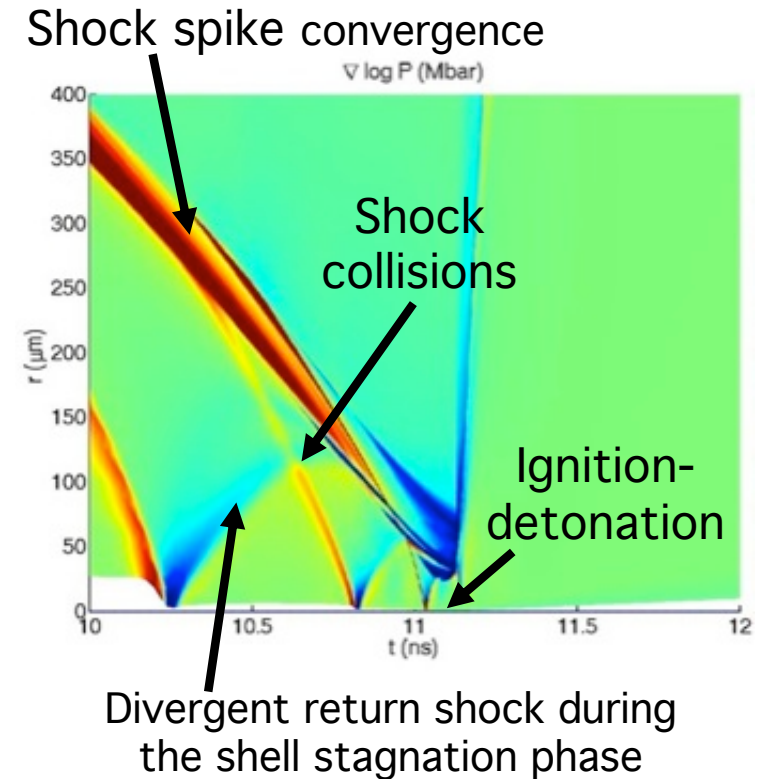
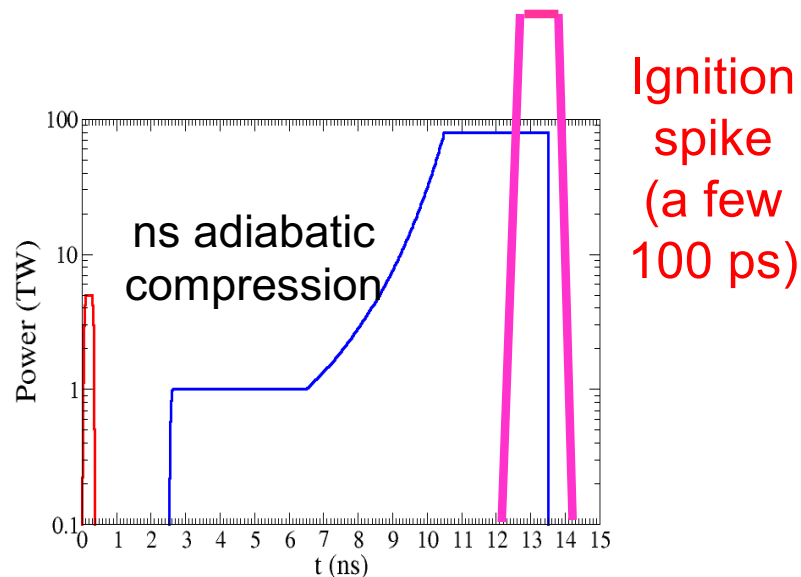
Decoupling compression and ignition phases

- Fast Ignition
- Shock Ignition



# Shock Ignition

- Scheme proposed by R. Betti, J.Perkins et al. [PRL 98 (2007)] and anticipated by V. A. Shcherbakov [Sov.J. Plasma Phys. 9, 240 (1983)]
- Thicker and more massive target  
Lower implosion velocity  $V \approx 240$  km/s

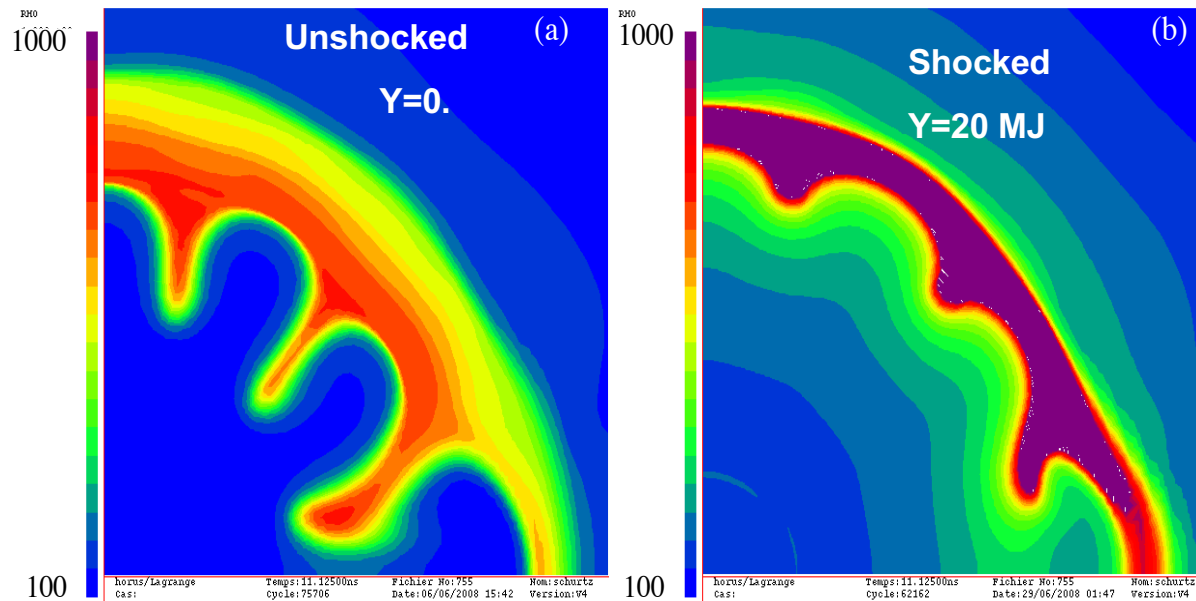
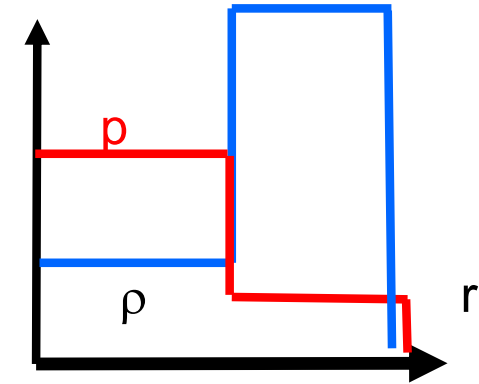


- A final laser spike launches a converging shock: at least 300 Mbar at the ablation front)
- The ignition shock collides with the return shock and provides the necessary amount of energy to trigger ignition from the central hot spot



# Advantages of Shock Ignition

- 1) Since compression phase does not provide a central hot spot, we can implode a thicker target at lower velocity, much less sensitive to hydro instabilities
- 2) Non isobaric fuel assembly implies higher gains



Simulations by S.Atzeni

In addition RT growth can also be mitigated due to competition between Rayleigh-Taylor and Richtmyer-Meshkov

Shock ignition is compatible with present-day laser technology 😊

# Unknowns of Shock Ignition



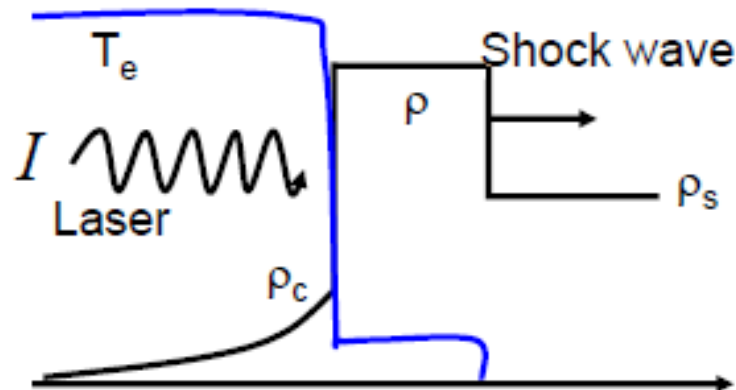
- Effect of **laser-plasma instabilities** at intensities up to  $\approx 10^{16}$  W/cm<sup>2</sup>. SRS, SBS and TPD. How they develop? How much light do they reflect?
- Are there many **hot electrons** and at what energy? What is their effect? *(usually in ICF hot electrons are dangerous since they preheat the target... Here they came at late times, large fuel  $\rho r$ , so they could indeed be not harmful or even beneficial, increasing laser-target coupling in presence of a very extended plasma corona...)*

For more information:

D. Batani, S. Baton, A. Casner, S. Depierreux, M. Hohenberger, O. Klimo, M. Koenig, C. Labaune, X. Ribeyre, C. Rousseaux, G. Schurtz, W. Theobald, V. T. Tikhonchuk  
 «[Physical issues in shock ignition](#)» Nuclear Fusion, 54 (2014) 054009

# Difference between classical ablation pressure and hot electron driven pressure

isothermal corona: laser ablation



$$I = 4\rho_c C_s^3$$

$$C_s \propto T_e^{1/2}$$

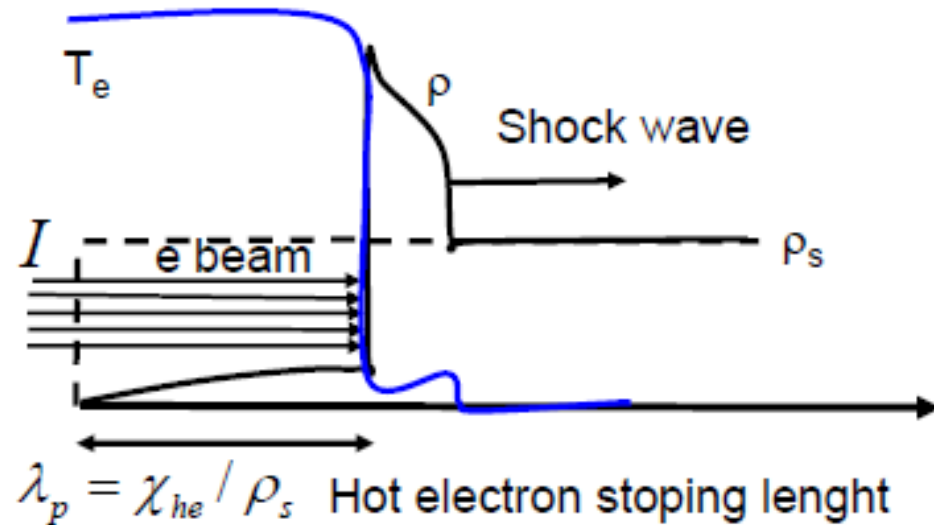
$$P \propto \rho_c C_s^2$$

$$T_e \propto \left(\frac{I}{\rho_c}\right)^{2/3}$$

$$P \propto I^{2/3} \rho_c^{1/3}$$

$\rho_c$  Critical density  
(0.03 g/cc at 0.35  $\mu\text{m}$ )

isothermal corona: hot electron ablation



$$\lambda_p = \chi_{he} / \rho_s \quad \text{Hot electron stopping length}$$

Time to establish pressure

$$t_s \propto \frac{\chi_{he}}{I^{1/3} \rho_s^{2/3}}$$

$$P \propto I^{2/3} \rho_s^{1/3}$$

$\rho_s$  Solid density  
(10 g/cc for SI DT)

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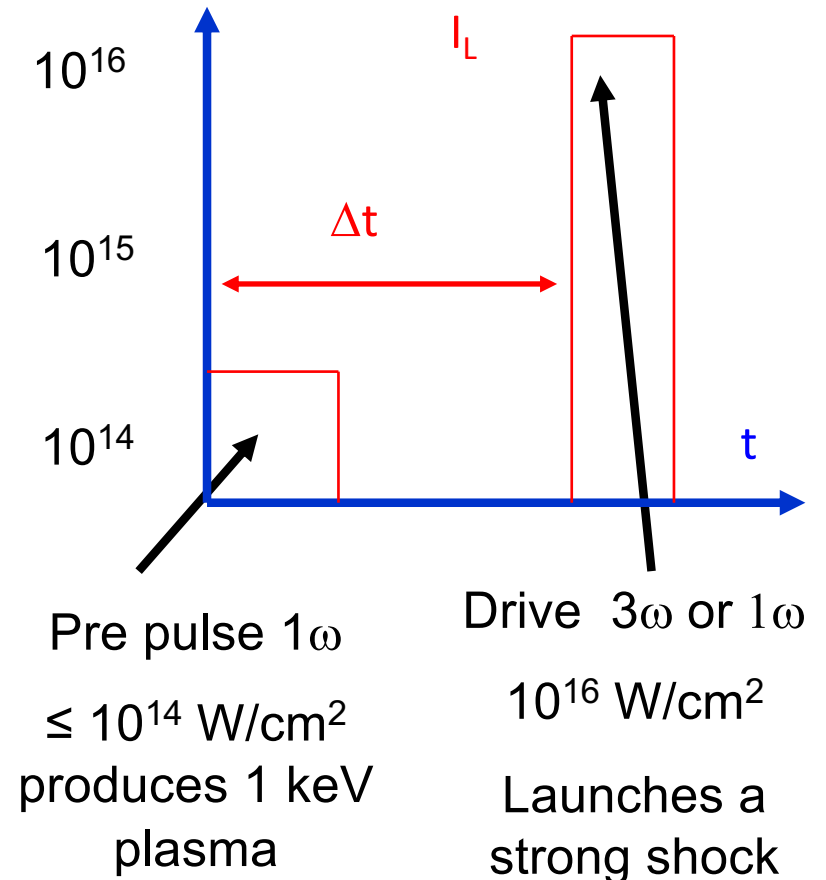
# Experiments at PALS - Prague



The PALS Iodine Laser

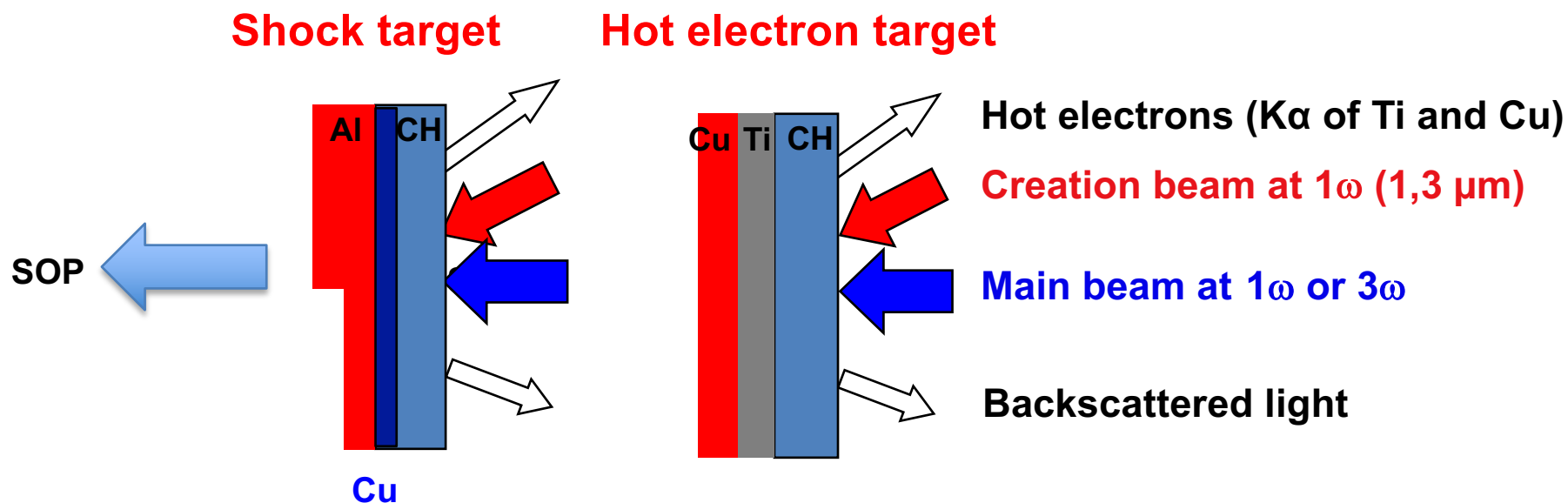
$\lambda = 1.3 \mu\text{m}$   $\tau = 300 \text{ ps}$   $E = 1500 \text{ J}$

$3\omega$   $\lambda = 0.44 \mu\text{m}$   $E \leq 500 \text{ J}$



D. Batani, L. Antonelli, V. Tikhonchuk, J. Badziak, K. Jakubowska, Z. Kalinowska, T. Pisarczyk, M. Rosinski, G. Cristoforetti, P. Koester, L.A. Gizzi, S. Atzeni, A. Schiavi, M. Skoric, S. Gus'kov, J. Honrubia, J. Limpouch, O. Klimo, O. Renner, M. Krus, J. Ullschmied et al. "Progress in understanding the role of hot electrons for the shock ignition approach to inertial confinement fusion" Nucl. Fusion 59 (2019) 032012

- ➔ one beam to create the plasma :  $1\omega$ , 300 ps - 50J,  $I \sim 10^{14}$  W/cm<sup>2</sup>, RPP Ø 300  $\mu$ m
- ➔ one beam to launch the shock :  $1\omega$  or  $3\omega$ , 300 ps - 500 J,  $I \sim 10^{16}$  W/cm<sup>2</sup>, Ø 100  $\mu$ m



The CH layer simulates the low-Z material of the shell of a pellet.

The Al layer is a standard material for shock measurements

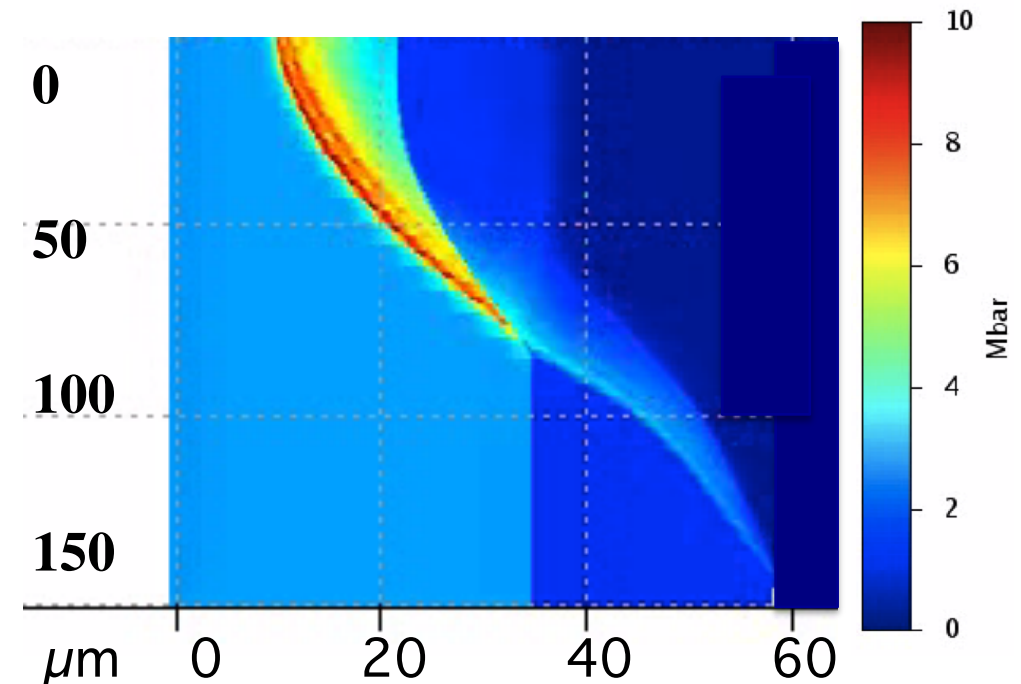
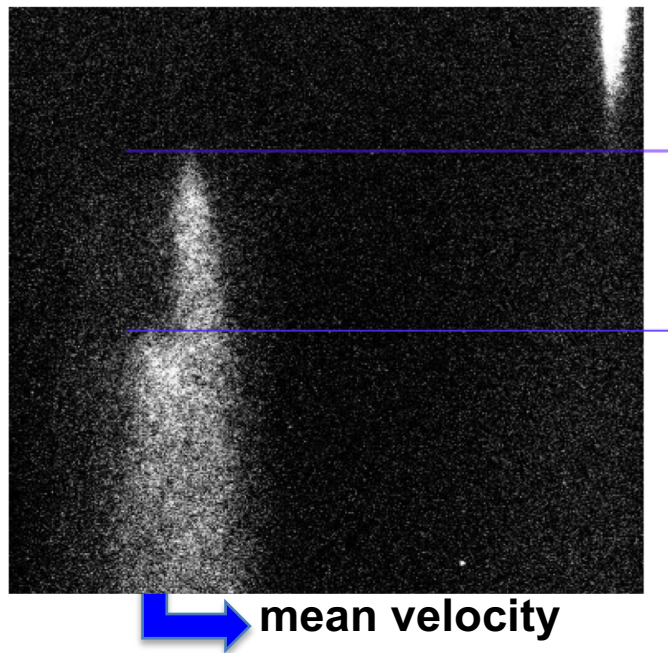
Cu and Ti used as tracer layer for  $K\alpha$  emission

# Shock chronometry for estimating the pressure (PALS)

Measured  $P$  at rear side much lower than ablation pressure at front side:  
Shock pressure undergoes a rapid decrease due to:

- 1) 2D effects during propagation
- 2) Relaxation waves from front side when laser turns off

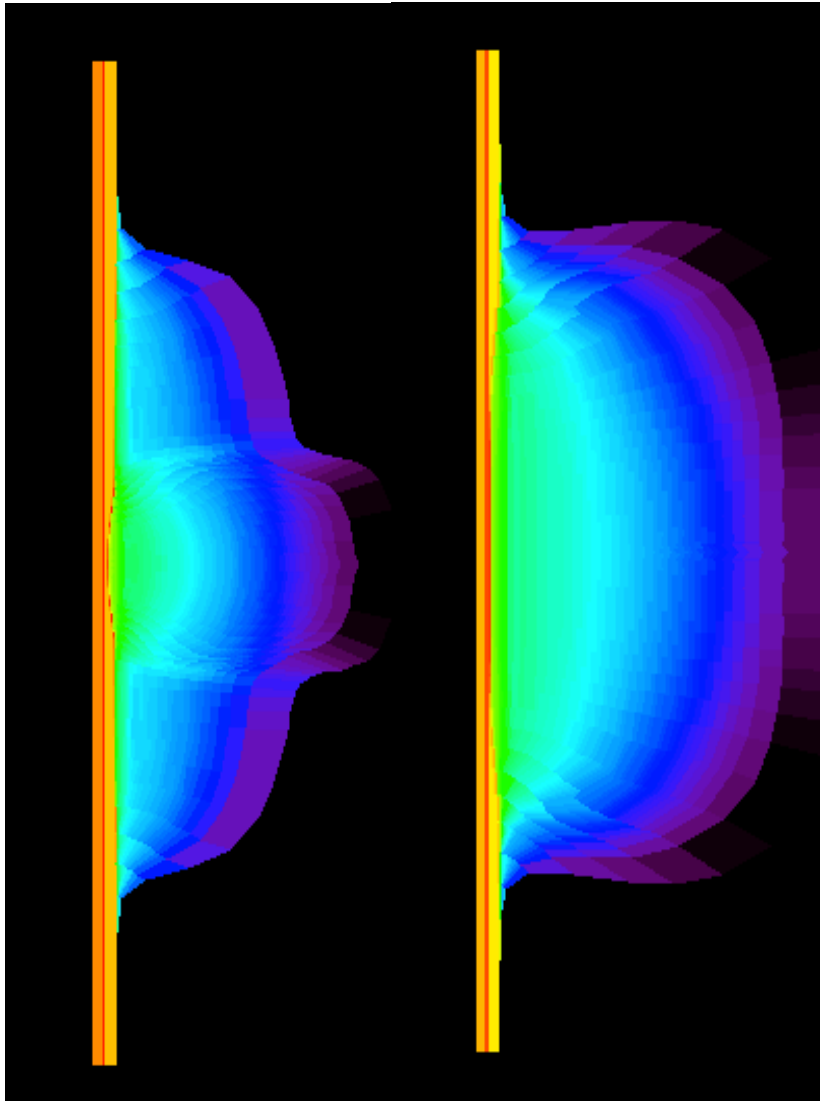
Target  
25  $\mu\text{m}$  CH  
35  $\mu\text{m}$  Al



We run hydro simulations to match shock breakout time and we find that a final pressure  $\leq 10$  Mbar corresponds to a maximum  $P \approx 90$  Mbar during interaction.

# 2D Hydro simulations

Initial ablation pressure  $\approx 90$  Mbar  $\ll$  estimation from scaling laws



## Explanations ?

*Lateral heat transport in the overdense region is important and reduces the shock pressure*

(In our experiment spot size comparable to the distance between critical layer and ablation surface  $\approx 40 \mu\text{m}$  vs.  $\approx 100 \mu\text{m}$ )

Simulations with the same laser parameters but larger spot ( $\geq 400 \mu\text{m}$ ) yield pressure  $\approx 180$  Mbar

Still simulations reproduce well the trend of data but cannot retrieve the expected pressure at  $10^{16} \text{ W/cm}^2$  (300 Mbar and not 180 ! )



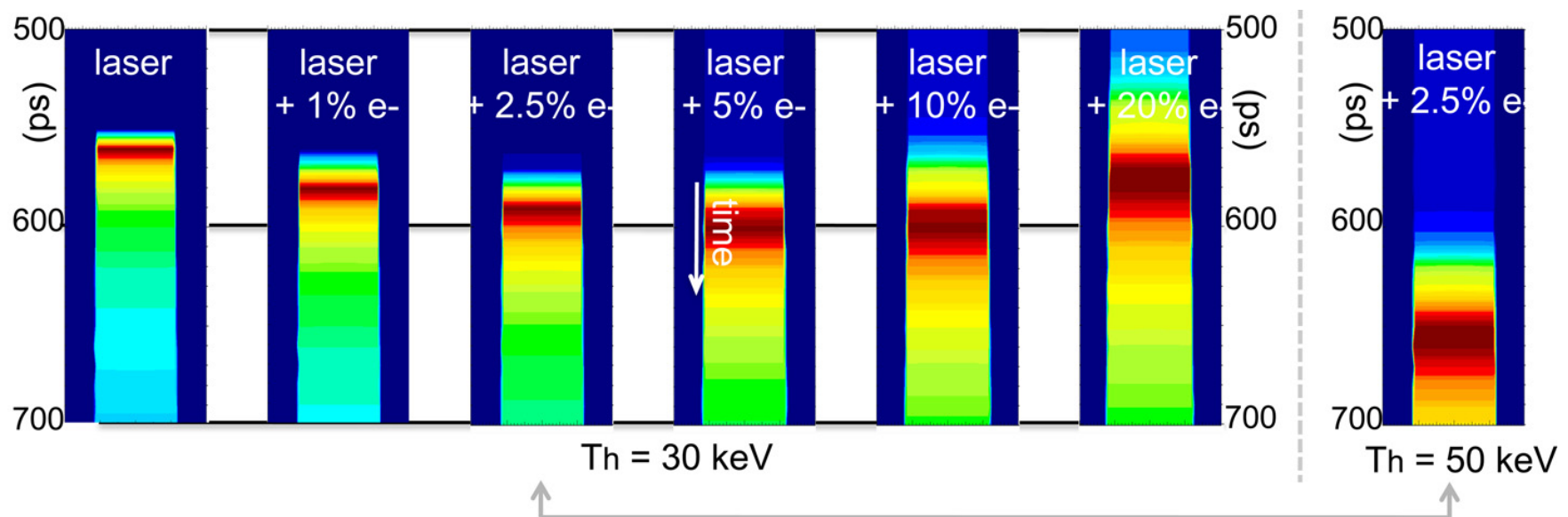
# What did we forget?

Hot electrons may preheat the target material and induce a “preheat pressure”  $P^*$ .

This results in:

- Decrease of compression
- Increase of shock velocity (early shock breakout)
- Expansion of target rear side (delayed shock breakout)

$$\rho^* = \rho_0 \frac{(\gamma + 1)P^* + (\gamma - 1)P_0}{(\gamma + 1)P_0 + (\gamma - 1)P^*} \quad D_s = \sqrt{\frac{(\gamma + 1)P^* + (\gamma - 1)P_0}{2\rho_0}}$$



# Simulations with “improved” model

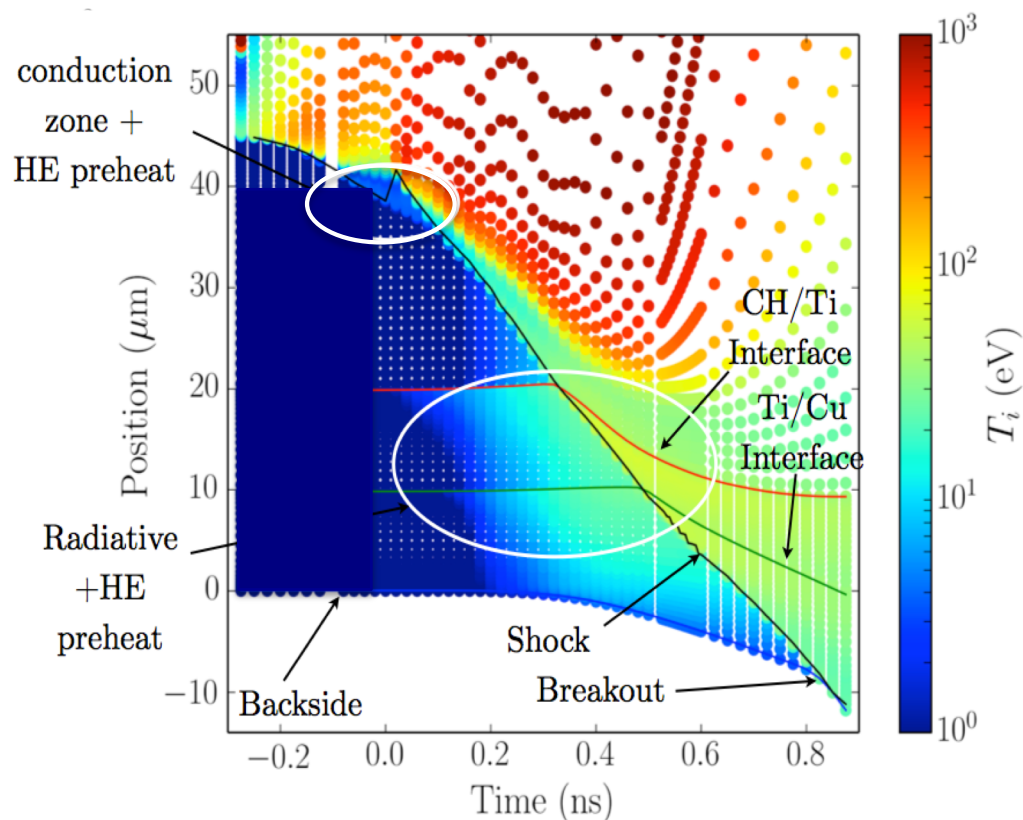
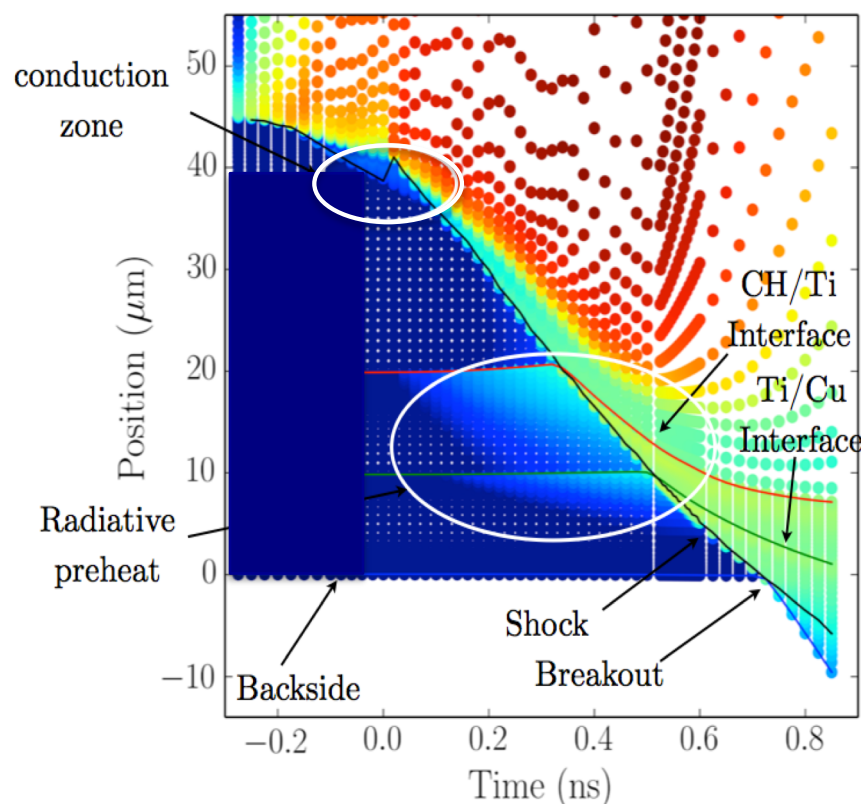
## “IMPROVED” HYDRO SIMULATION CODES:

- Better description of absorption (PCGO: from ray tracing to gaussian beamlets)
- Real time treatment of parametric instabilities and resonant absorption
- Generation of hot electrons and coupling to hydro (simplified kinetic transport)

### CH/Ti/Cu target

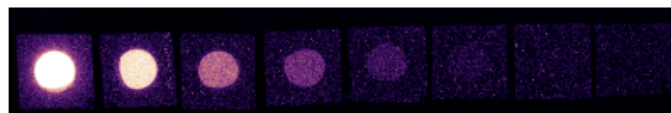
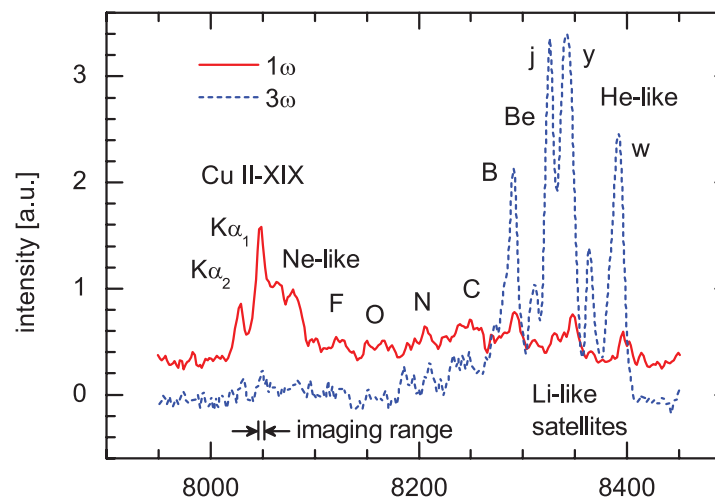
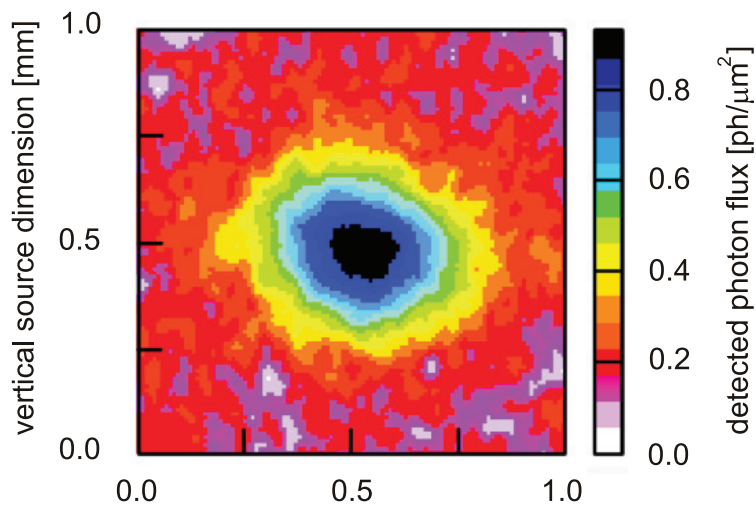
### Without hot electrons

### With hot electrons

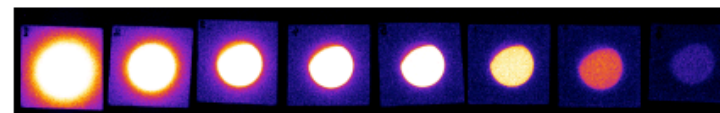


“Coupled hydrodynamic model for laser-plasma interaction and hot electron generation” A. Colaitis, et al. PHYSICAL REVIEW E 92, 041101(R) (2015)

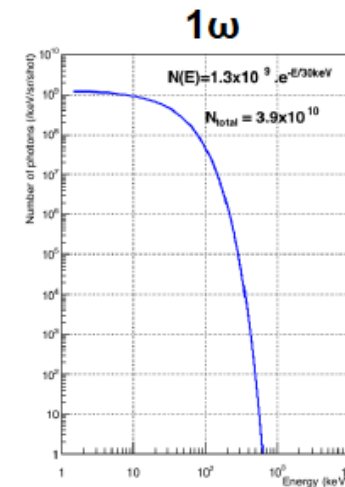
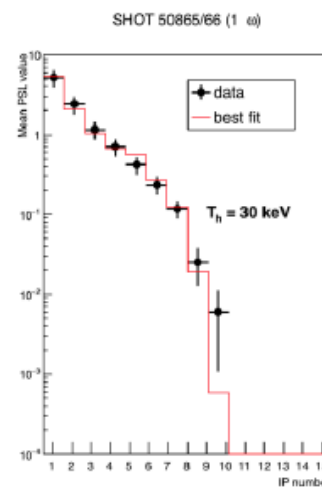
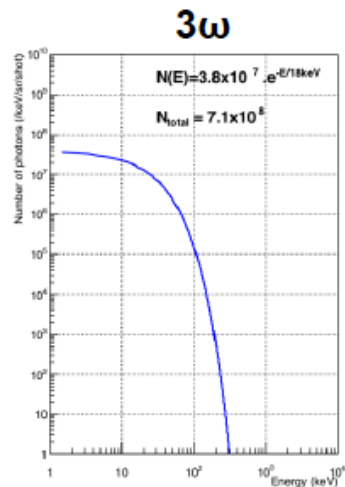
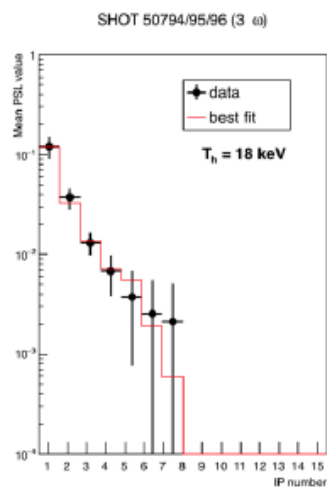
Simultaneous characterisation by: i) X-ray K- $\alpha$  imaging, ii) X-ray spectroscopy, iii) Bremsstrahlung spectroscopy



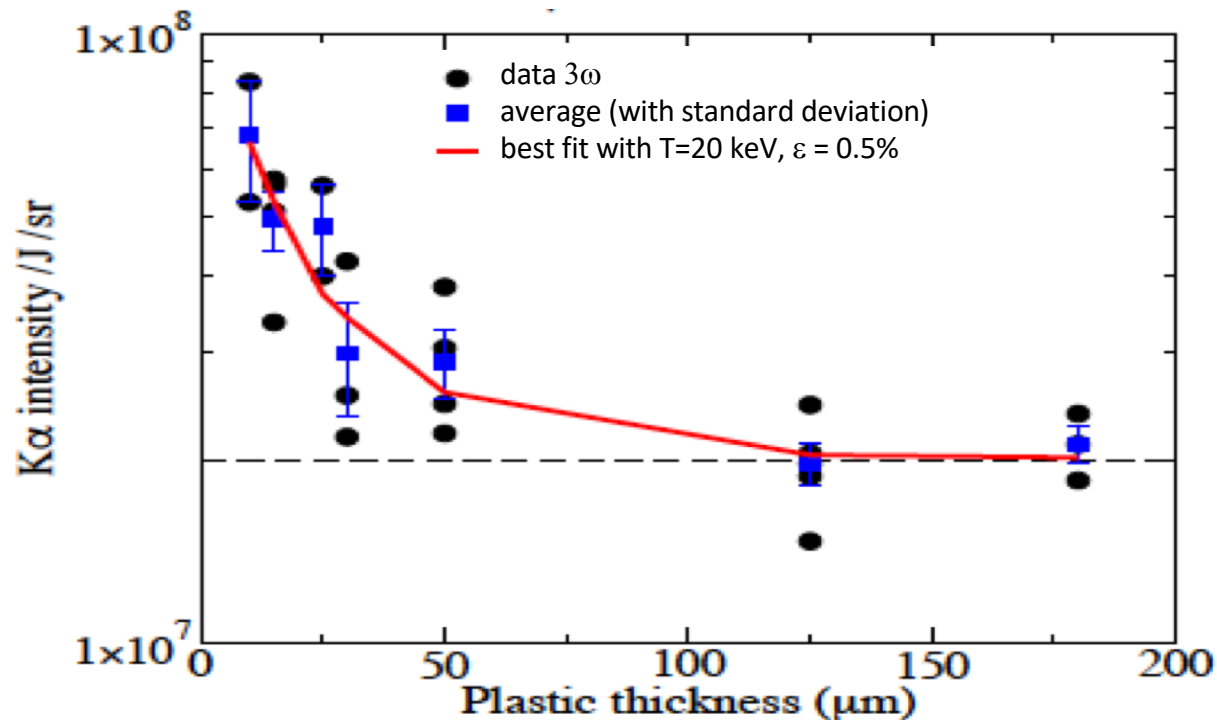
SHOT 50794/95/96 (3  $\omega$ )



SHOT 50865/66 (1  $\omega$ )



## Hot electrons: results



Ti Photon fluxes reconstructed from x-ray spectroscopy for  $3\omega$  irradiation

**Main source of hot electrons in our conditions is SRS**

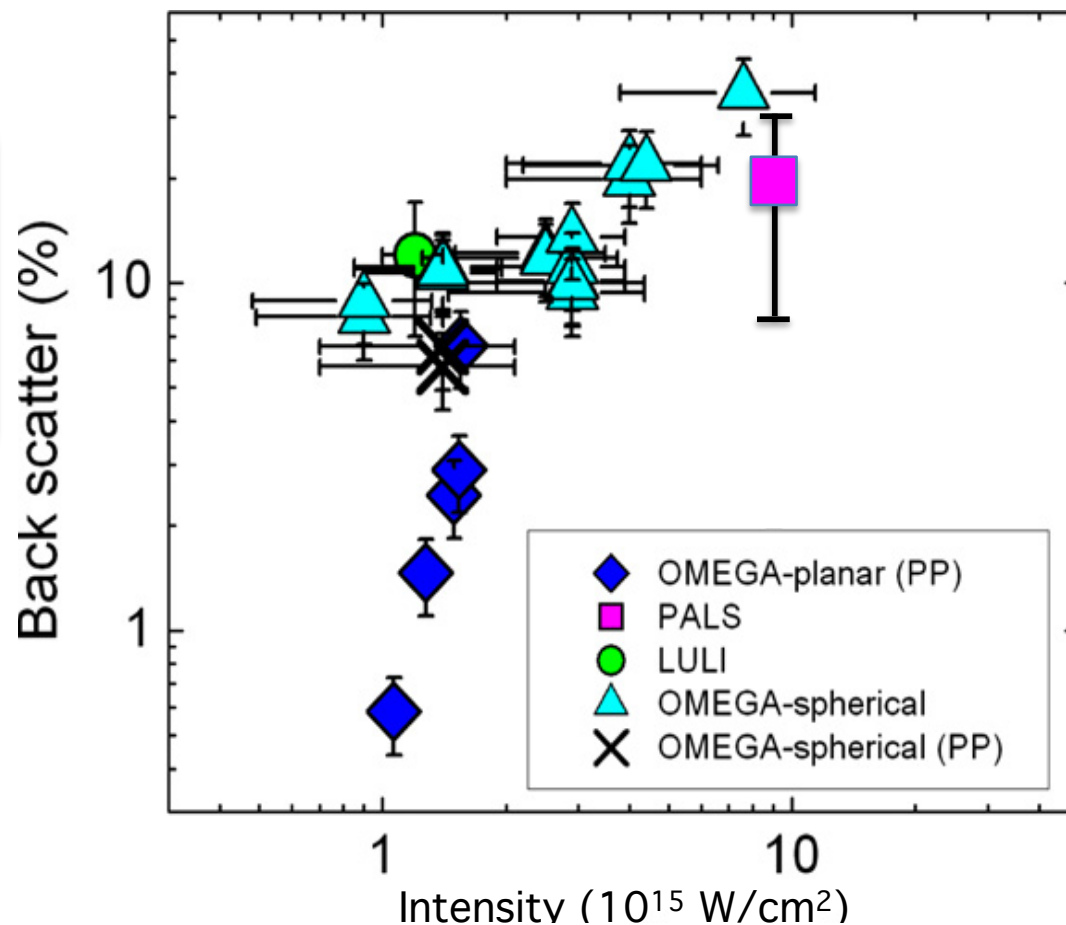
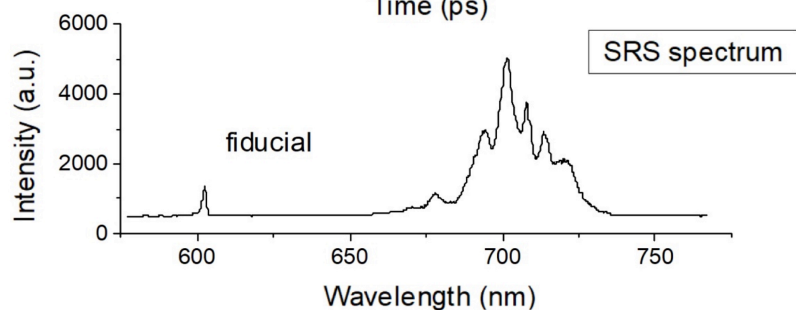
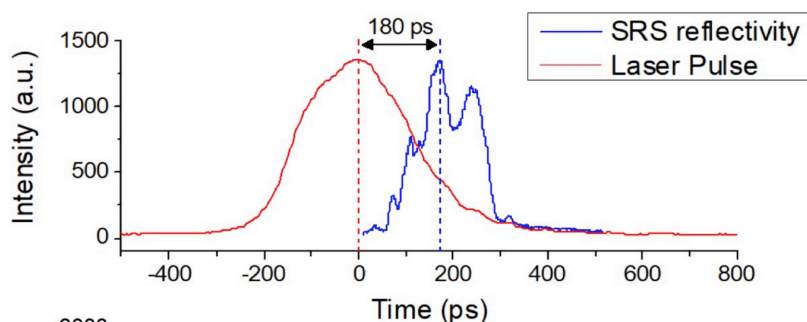
HE characteristics at  $3\omega$  and  $1\omega$

	$3\omega$ , $E_{\text{laser}} \approx 200\text{J}$	$1\omega$ , $E_{\text{laser}} \approx 650\text{J}$
$T_{e^-}$ (keV)	$20^{+15}_{-8}$	$38^{+57}_{-12}$
$\varepsilon_{\text{laser} \rightarrow e^-}$ (%)	$0.65^{+1}_{-0.14}$	$2.66^{+3.45}_{-0.13}$

# Parametric Instabilities

## Reflectivity

- $\leq 10\%$  inside lens cone
- $\leq 15\%$  outside
- mainly dominated by SRS



# PALS expt: conclusions

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- At  $3\omega$ , hot electron temperature of  $\sim 20$  keV and an energy conversion  $\leq 1\%$ , have been measured
- Reflectivity  $\leq 25\%$  mainly dominated by SBS
- At  $1\omega$ , both hot electron temperature and energy conversion increase ( $T \sim 40$  keV,  $\varepsilon \leq 3\%$ )
- Main source of hot electrons in our conditions is SRS
- Lateral heat transport in the overdense region is important and reduces the shock pressure
- CHIC simulations reproduced the hot electron characteristics at  $3\omega$  and (with some modifications to the scaling laws) at  $1\omega$
- In order to reproduce experimental results it is essential to take the effects of hot electrons into account self-consistently.

# Summary

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- Omega EP

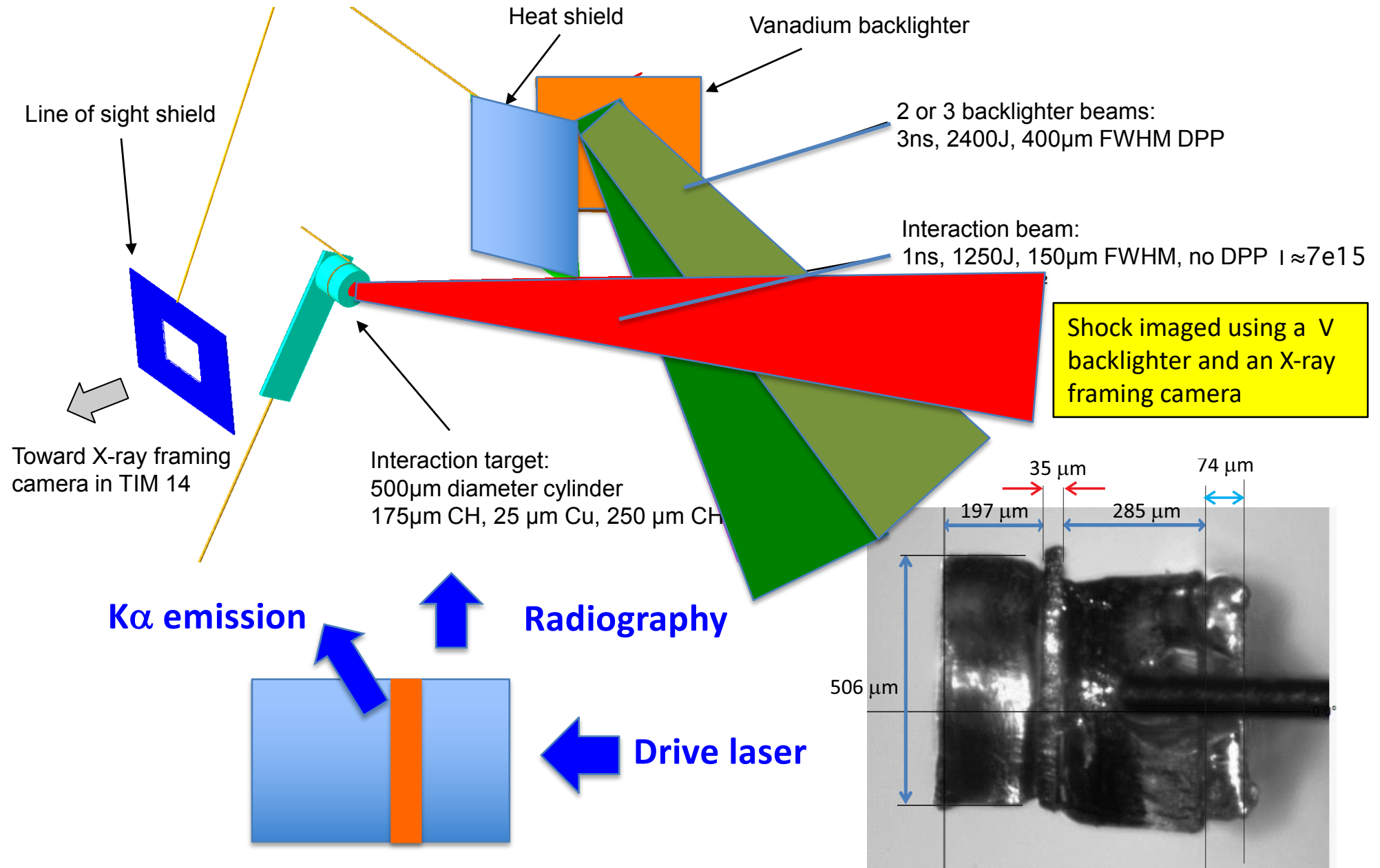
- Preliminary results from planar experiments on formation of strong shocks and hot electron effects on shock dynamics

- Experiment supported by EUROfusion ...

- Collaboration experiment involving CELIA, JIHT, UPM, INO-CNR, UOY, RAL, LULI URome, UCanaria

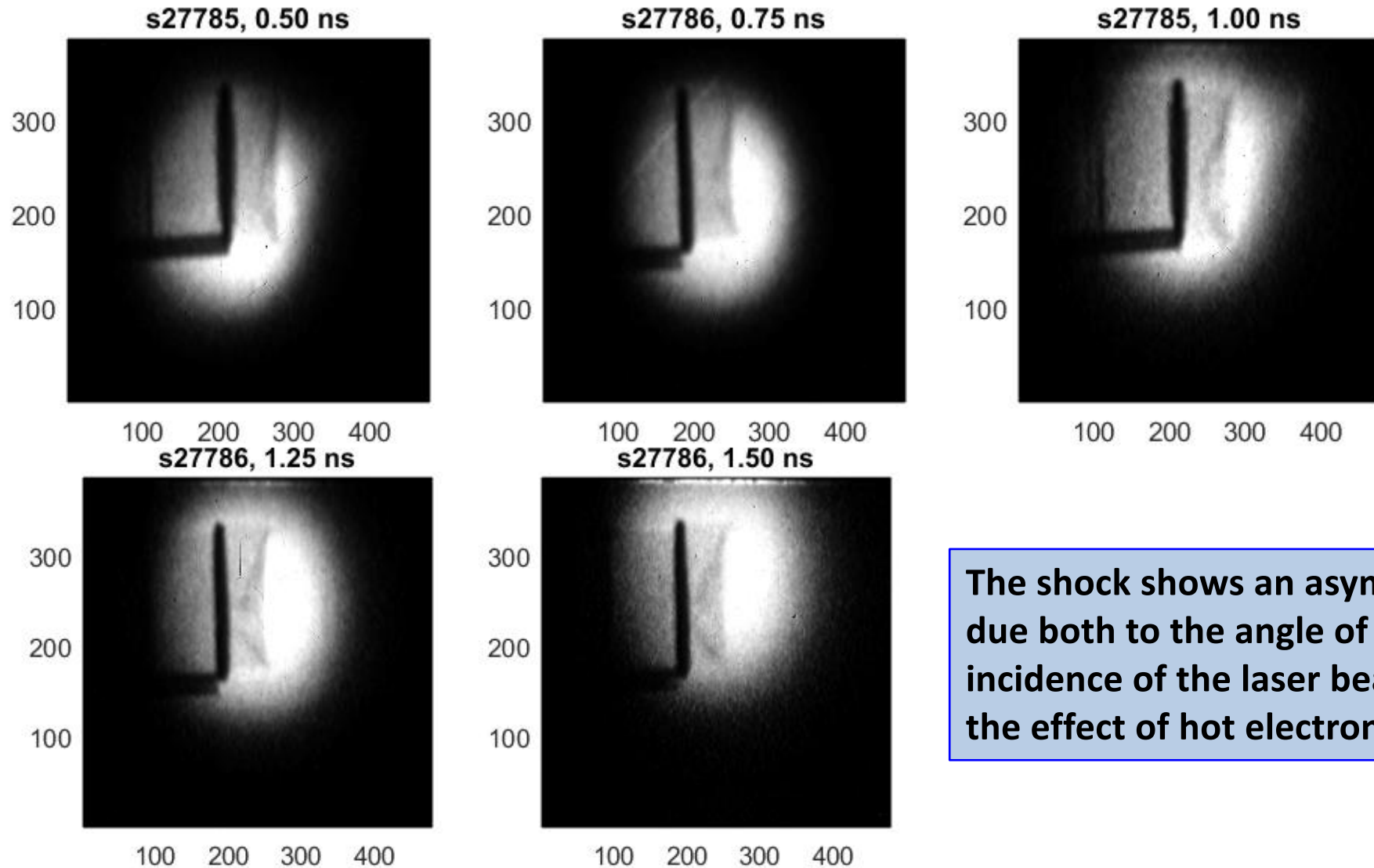


# Experiment done on OmegaEP to evaluate effect of hot electrons on shock hydrodynamic (planar geometry)



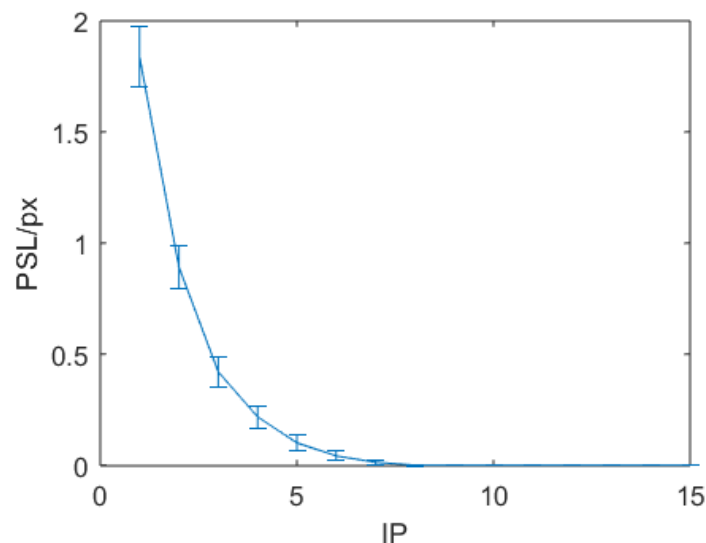
# March Shot Day: 2D time-resolved radiographs

radiography images at several times allows to reconstruct the shock propagation inside the front CH.



The shock shows an asymmetry due both to the angle of incidence of the laser beam and the effect of hot electrons.

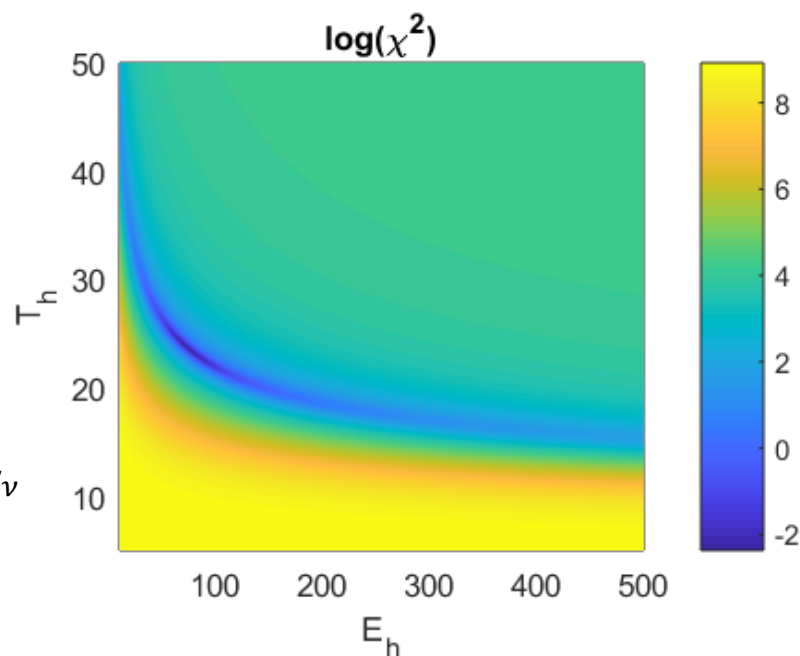
# Hot electron temperature and energy are measured using BMXS



Signal on IP of BMXS for shot 27785

- IP 8 to 15: see nothing.
- IP 1 to 4: see two temperature distribution ?
- IP 5 to 7: see hot electrons distribution !

For IP 5 to 7



$\chi^2$  is used to determine both HE temperature and energy:

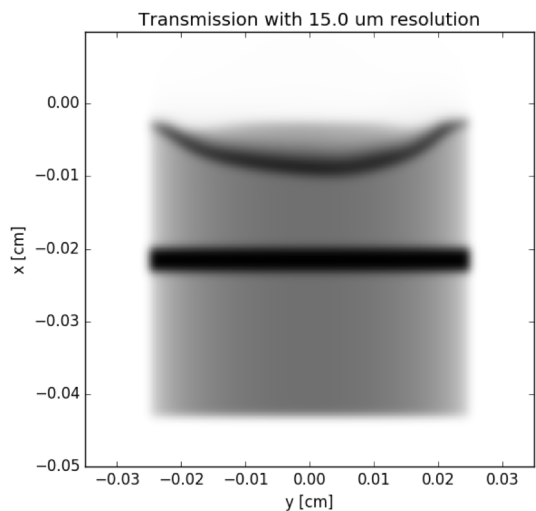
$$S_i(T_h, E_h) = \alpha_i \frac{E_h}{1.5 T_h} \int F_i(\varepsilon_\nu) \left[ \int I_{br}(\varepsilon_h, \varepsilon_\nu) \times f^M(\varepsilon_h, T_h) d\varepsilon_h \right] d\varepsilon_\nu$$

$$\tilde{\chi}^2(T_h, E_h) = \frac{1}{n-2} \sum_{i=1}^n \frac{[S_i(T_h, E_h) - Q_i]^2}{\delta S_i(T_h, E_h)^2 + \delta Q_i^2}$$

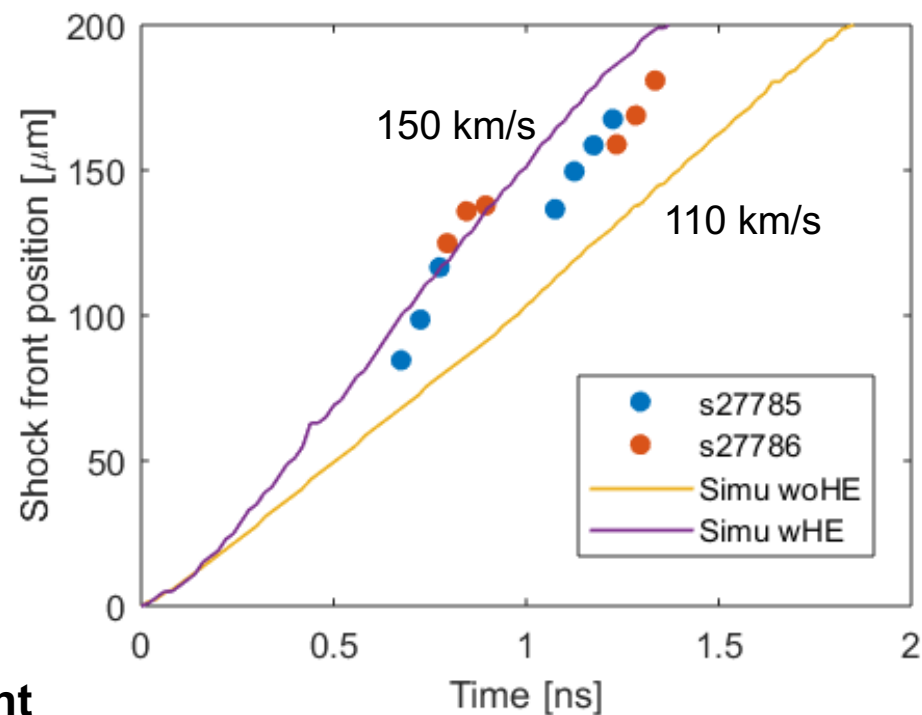
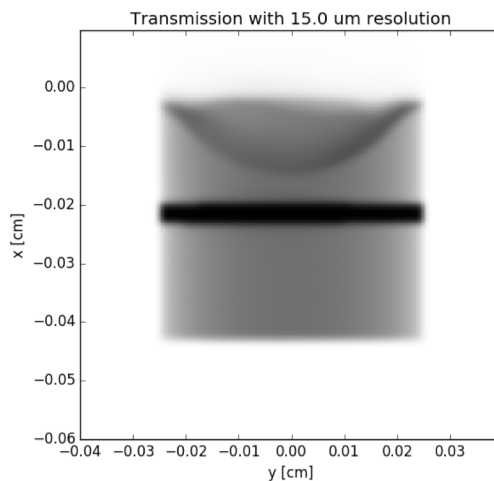
**Hot electrons measurement with BMXS show a HE temperature of 25 keV and an energy conversion of 5% *compatible with CHIC predictions***

# Synthetic radiographies with hot electrons show better agreement with experiment

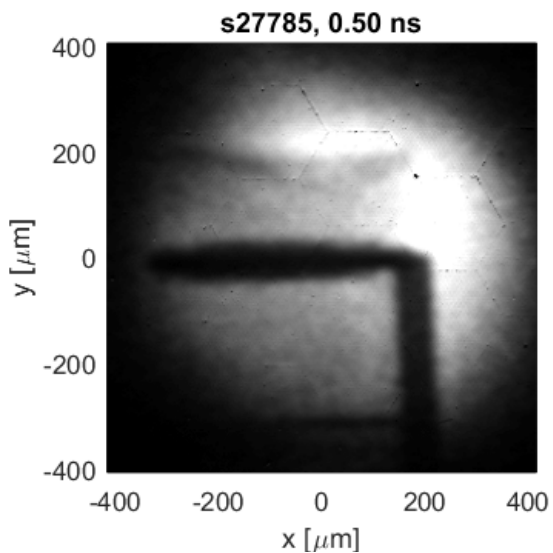
**Simulation woHE**



**Simulation wHE**

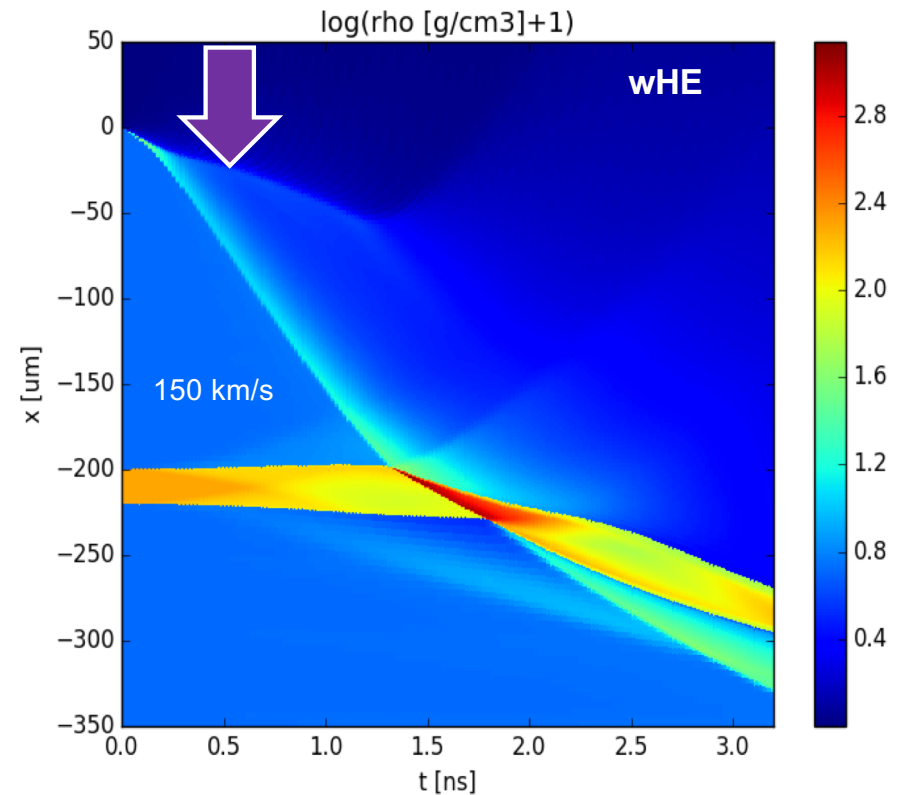
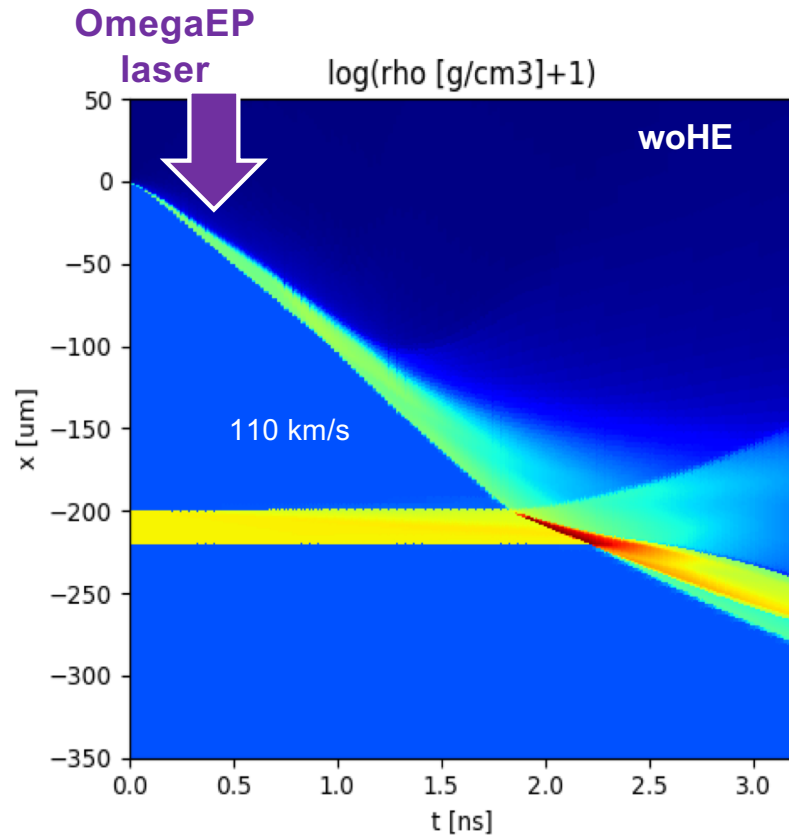


**0.5 ns**



**Experiment**

# Displaying the density in the (x,t) plan shows the effect of hot electron heating on shock dynamics

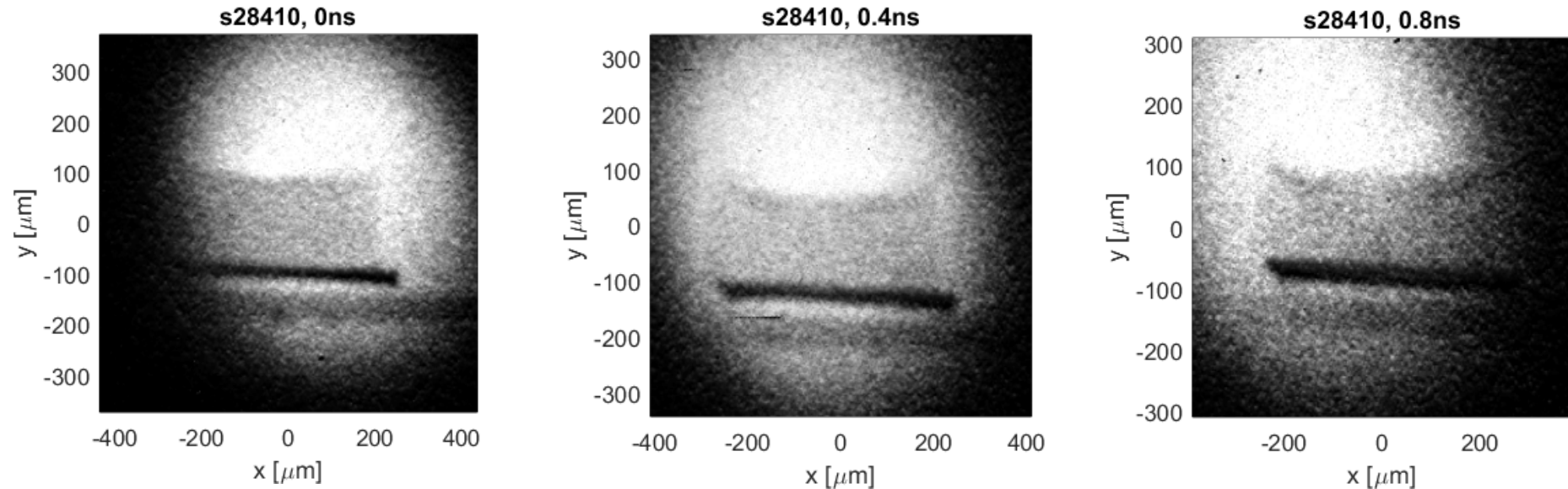


	woHE	wHE
Shock pressure in front CH	~100 Mbar	~150 Mbar
Shock breakout CH to Cu	1.82 ns	1.32 ns
Shock breakout rear of Cu	2.25 ns	1.82 ns

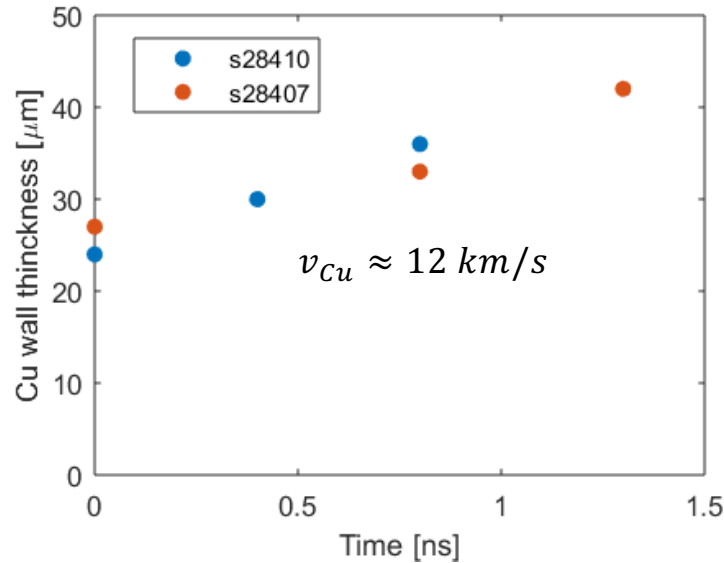
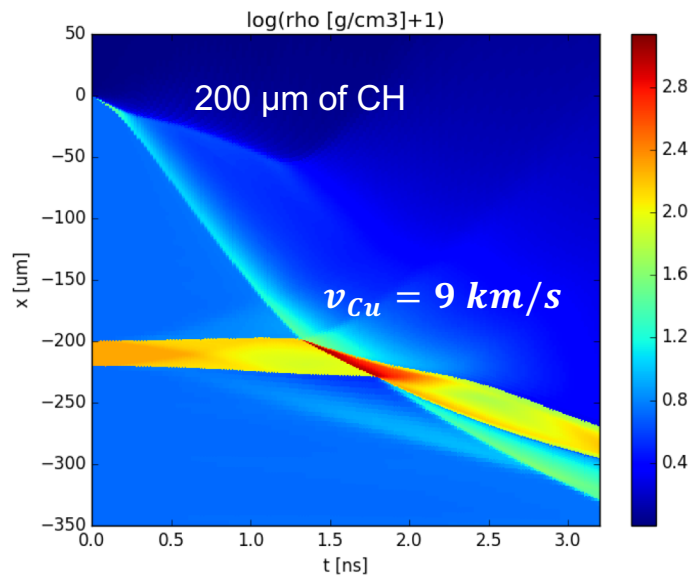
## Effect of the HE on the hydrodynamic:

- **Faster main shock**
- **Expansion of shocked CH**
- **Expansion of the Cu layer**

# The expansion of copper layer due to hot electron heating has been clearly observed.



175  $\mu\text{m}$  CH facing laser

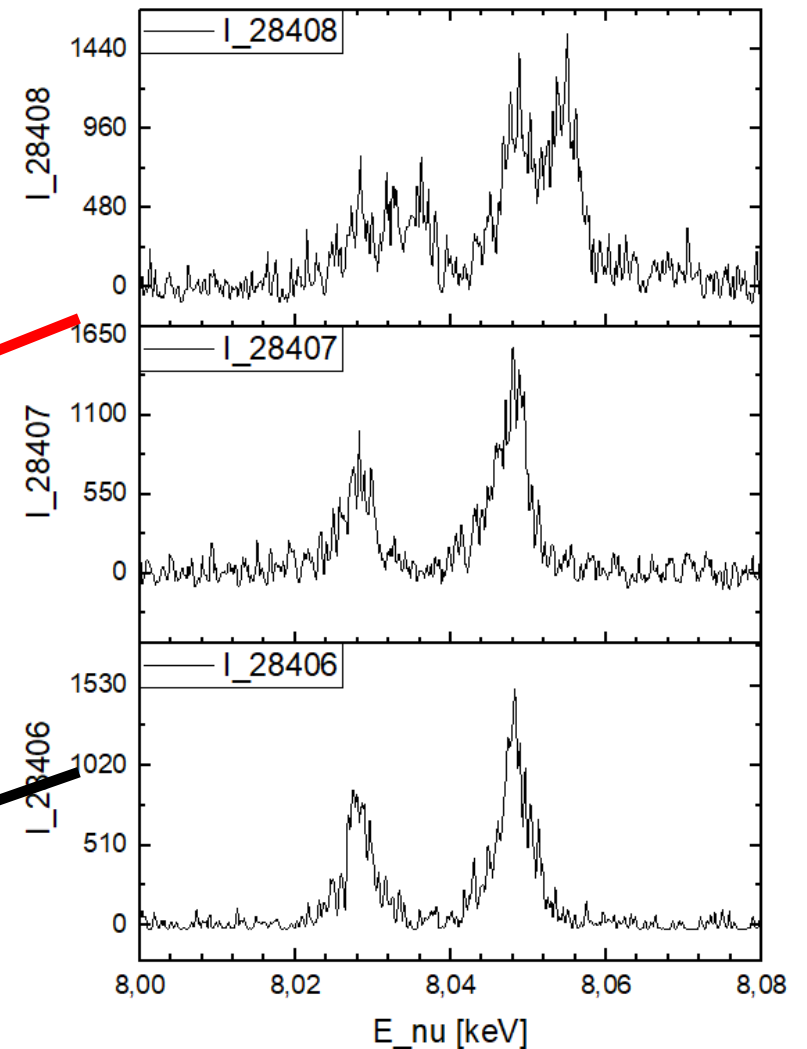


The measured expansion velocity of the copper layer is coherent with simulation

# $K_\alpha$ spectrometry

- Two diagnostics measure the  $K_\alpha$  emission from copper:
  - ZNVH: Zinc Von Hamos, crystal reflection recorded on IP.
  - HRS: High Resolution Spectrometer, recorder on CCD.

HRS raw spectra near Cu  $K_\alpha$  line



Broadened/split lines  
= WDM

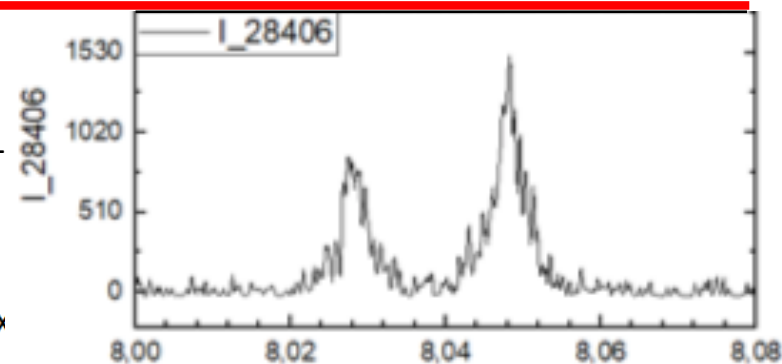
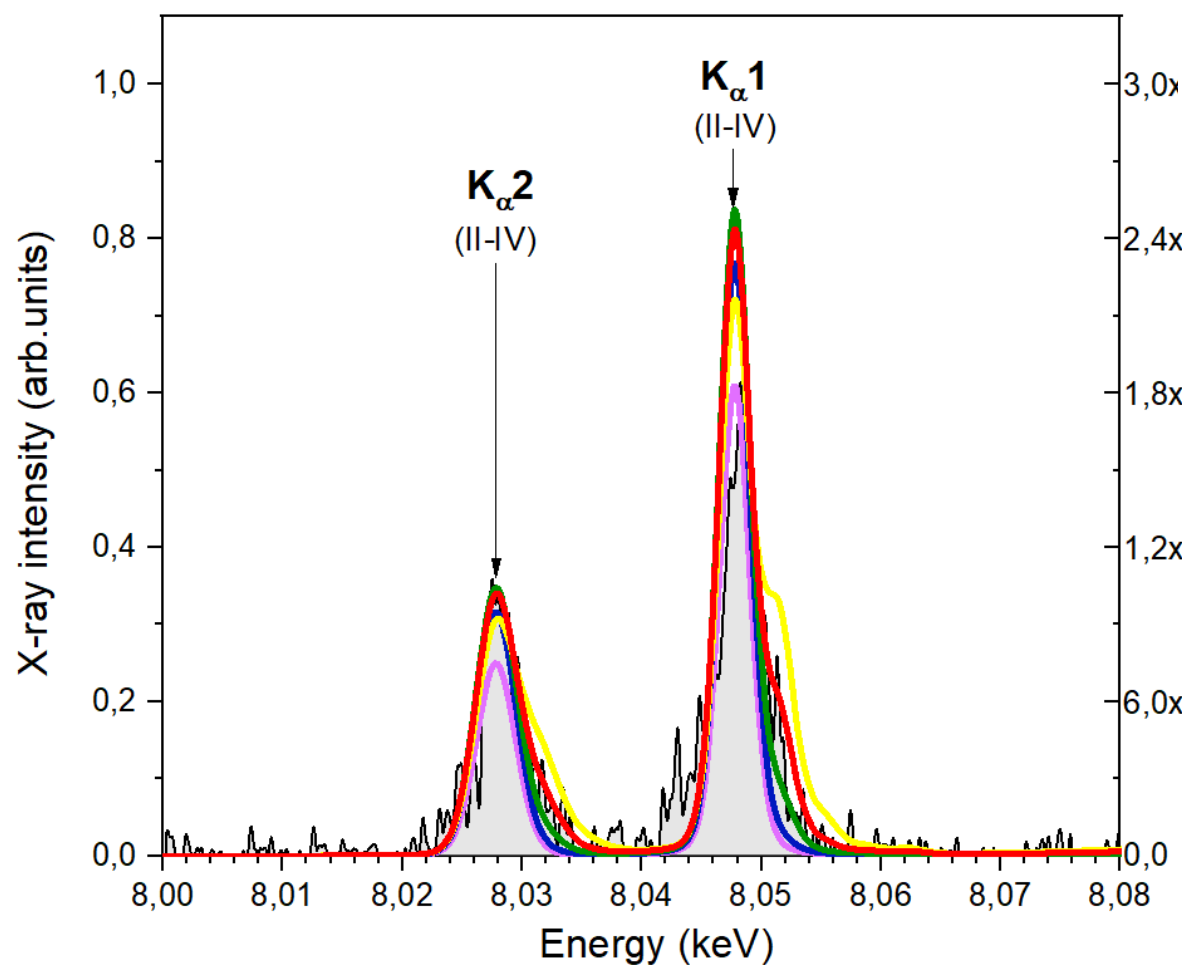
Changing beam energy and plastic overlayer thickness

Simple / narrow line case

# Modeling by PrismSpect

## Narrow line case

$Ni=8e22=N_{solid}$ ,  $Ne \sim 3e23 \text{ cm}^{-3}$ ,  $T_{hot}=10 \text{ keV}$  (1)



Emissivity (erg/cm<sup>3</sup>/ster/s/eV)

- Experiment I\_28406
- $T_e = 10 \text{ eV}$ , 1% hot 10 keV
- $T_e = 15 \text{ eV}$ , 1% hot 10 keV
- $T_e = 20 \text{ eV}$ , 1% hot 10 keV
- $T_e = 25 \text{ eV}$ , 1% hot 10 keV
- $T_e = 30 \text{ eV}$ , 1% hot 10 keV

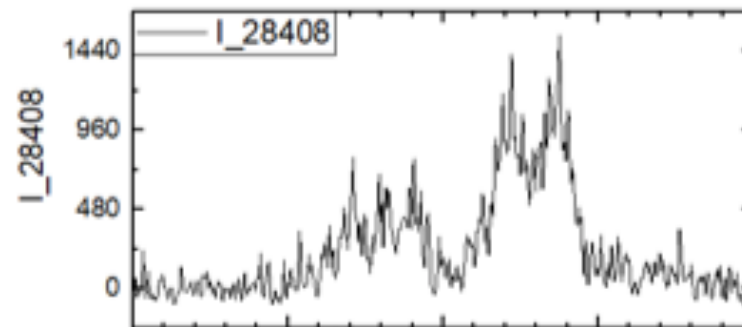
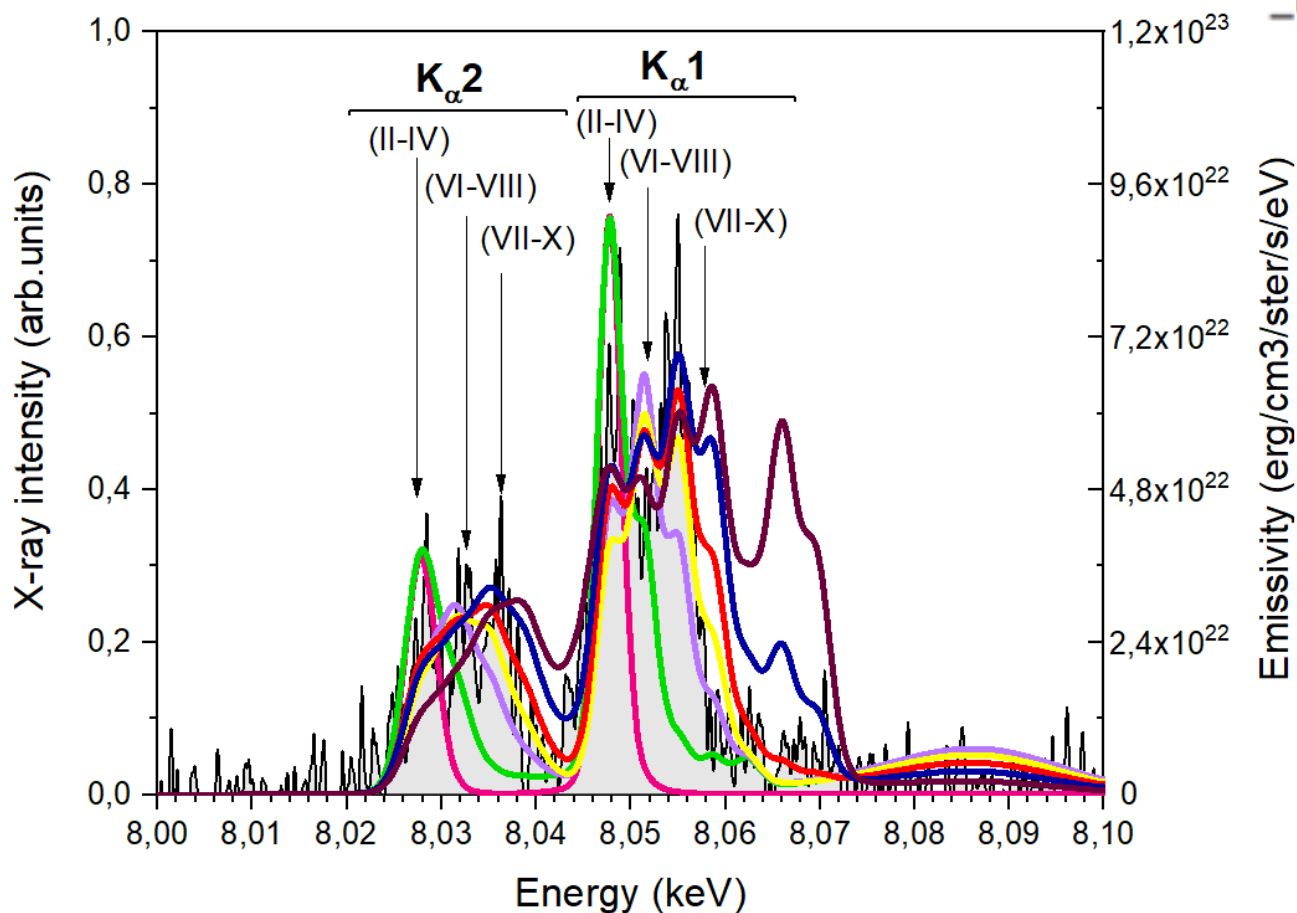
**Shot #28406:**  
 **$T_e = 22(\pm 5) \text{ eV}$**



# Modeling by PrismSpect

## Broad lines case

$T_{\text{hot}}=10$  keV (5%), Ni =  $8e22 = N_{\text{solid}}$ , Ne  $\sim 6e23$  cm $^{-3}$



- Experiment I\_28408
- $T_e = 10$  eV, 5% hot 10 keV
- $T_e = 30$  eV, 5% hot 10 keV
- $T_e = 50$  eV, 5% hot 10 keV
- $T_e = 70$  eV, 5% hot 10 keV
- $T_e = 90$  eV, 5% hot 10 keV
- $T_e = 110$  eV, 5% hot 10 keV
- $T_e = 130$  eV, 5% hot 10 keV

**Shot #28408**  
 **$T_e \sim 70(\pm 10)$  eV**

with definite contribution from colder "surrounding" target area ( $T_e < 20$  eV)

# Omega EP expt: conclusions

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- **Hot electron temperature of  $\sim 25$  keV and an energy conversion of 5%, i.e. a total energy of  $\sim 60$  J have been simultaneously measured, in agreement with predictions from CHIC simulations.**
- **CHIC simulations reproduced the shock velocity when the effects of hot electrons are taken into account. They predict an increase of 50% in shock pressure with hot electrons**
- **Synthetic radiographies roughly reproduce the shape of the shock front when hot electrons are taken into account**
- **The copper layer expands due to preheating induced by hot electrons and a WDM state is produced which is in agreement with results from  $K\alpha$  spectroscopy**

Both PALS and Omega EP experiments show that:

- We need to develop and optimize diagnostics dedicated to study the dynamics of very strong shocks in materials preheated by hot electrons and for the characterization of hot electrons
- Modelling of laser-plasma interaction and shock dynamics must self-consistently take into account the effects of parametric instabilities, the generation of hot electrons and the effect they have on target (increase in pressure but also preheating-induced expansion). Such modelisation requires developing and “tuning” advanced hydrodynamics codes

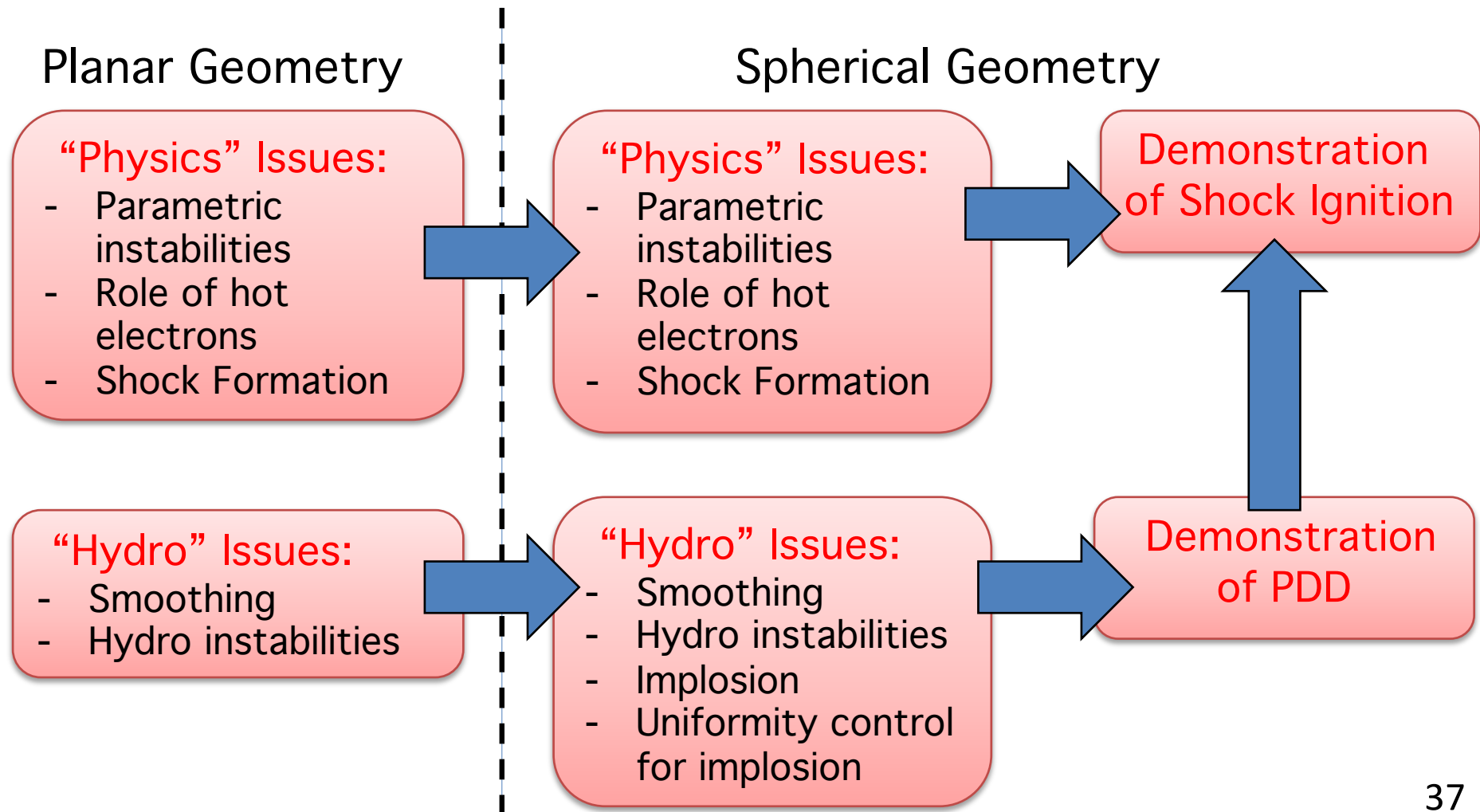
# Summary

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- 1) What is shock ignition
- 2) The problem of hot electrons and  
Experiments at PALS and Omega-EP
- 3) Roadmap to shock ignition

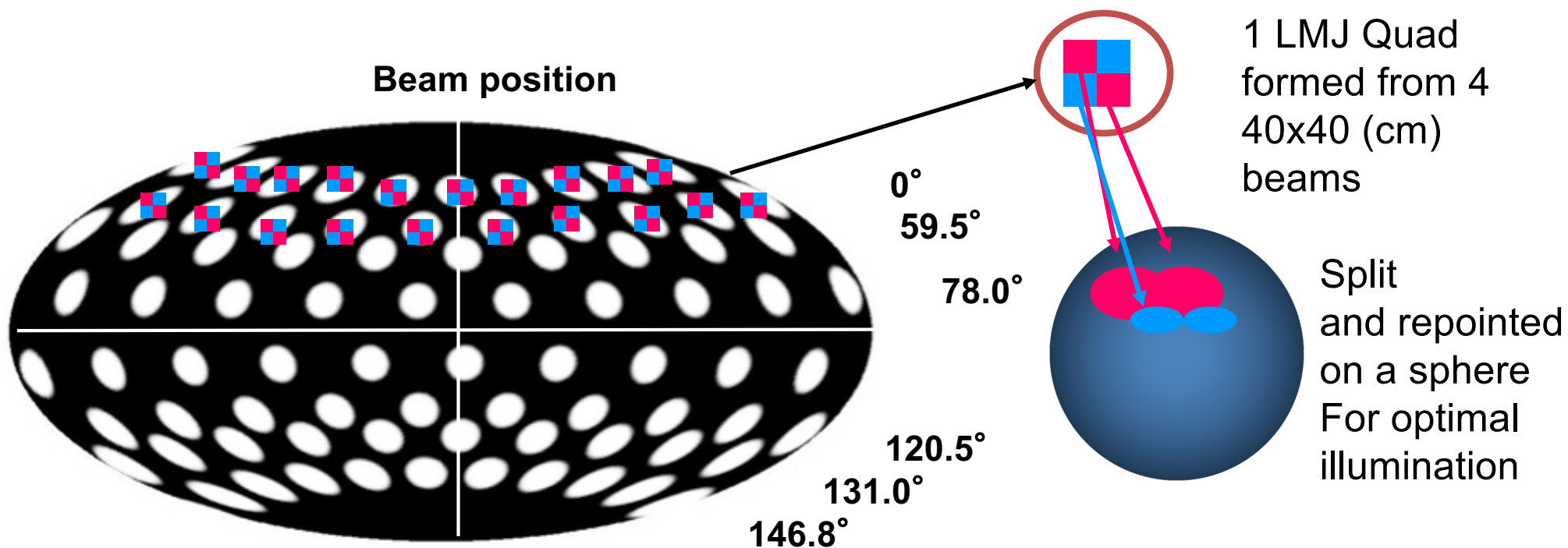
# Road map to shock ignition

How to approach the final goal of “Performing shock ignition demonstration experiments ” ?



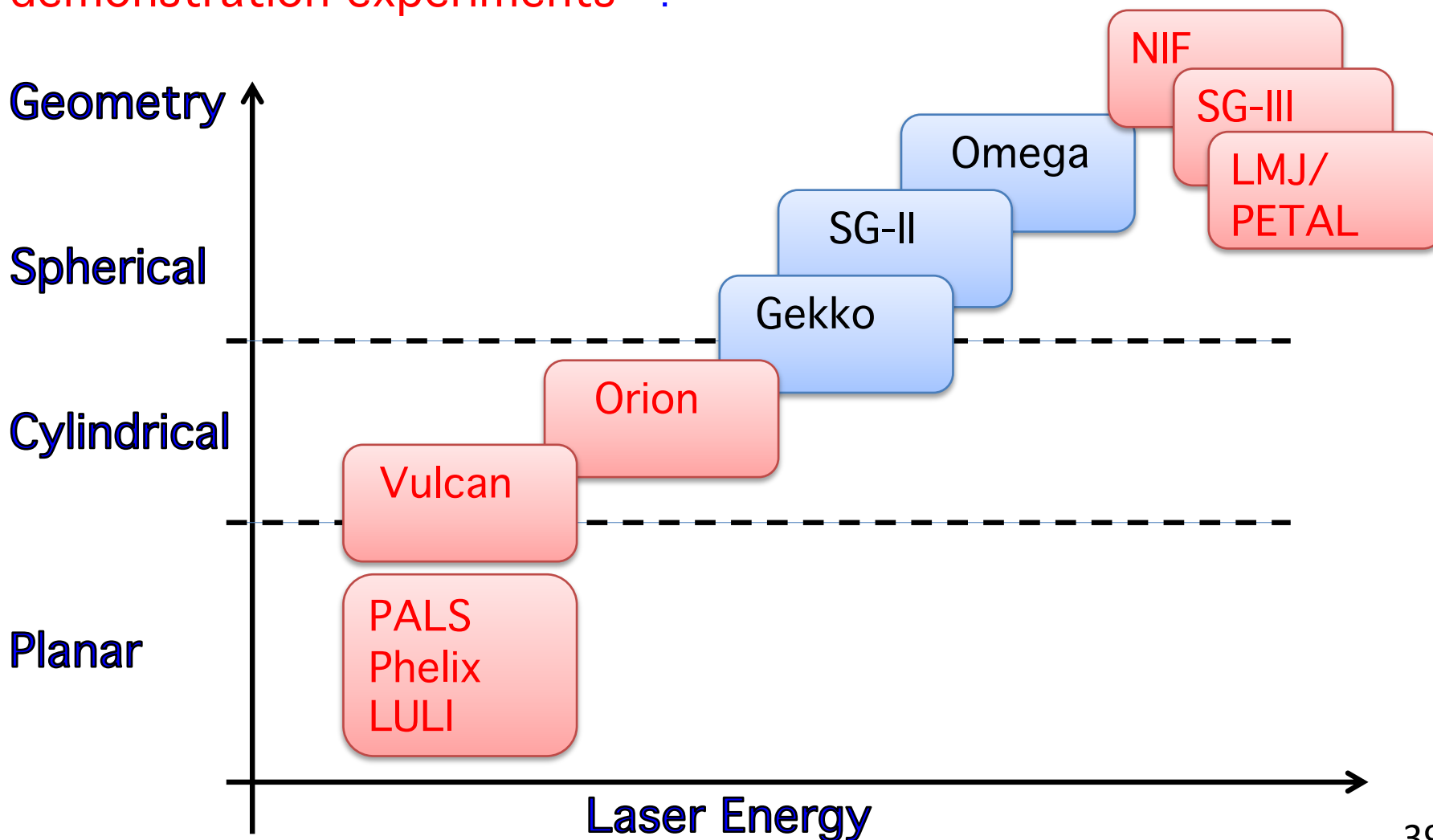
# Polar Direct Drive (PDD)

*An example of how it could be done...*



# Road map to shock ignition

How to approach the final goal of “Performing shock ignition demonstration experiments” ?



# Summary / Conclusions

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- Interesting physics which needs to be understood and mastered
- SI can be demonstrated at NIF, LMJ or or the Shenguang-III laser facility in China in the next decade
- BUT development of a full programme relies on:
  - Scientific credibility: physics issues addressed using:
    - ◆ “Smaller” facilities: PALS, ORION, Vulcan, LULI, Phelix, Gekko, Shenguang-II
    - ◆ intermediate-scale facilities: OMEGA, Shenguang-III P
- Support of EUROfusion has been, and is, important



**Thank you for your attention !**

